

Particle Detectors

Irakli Keshelashvili

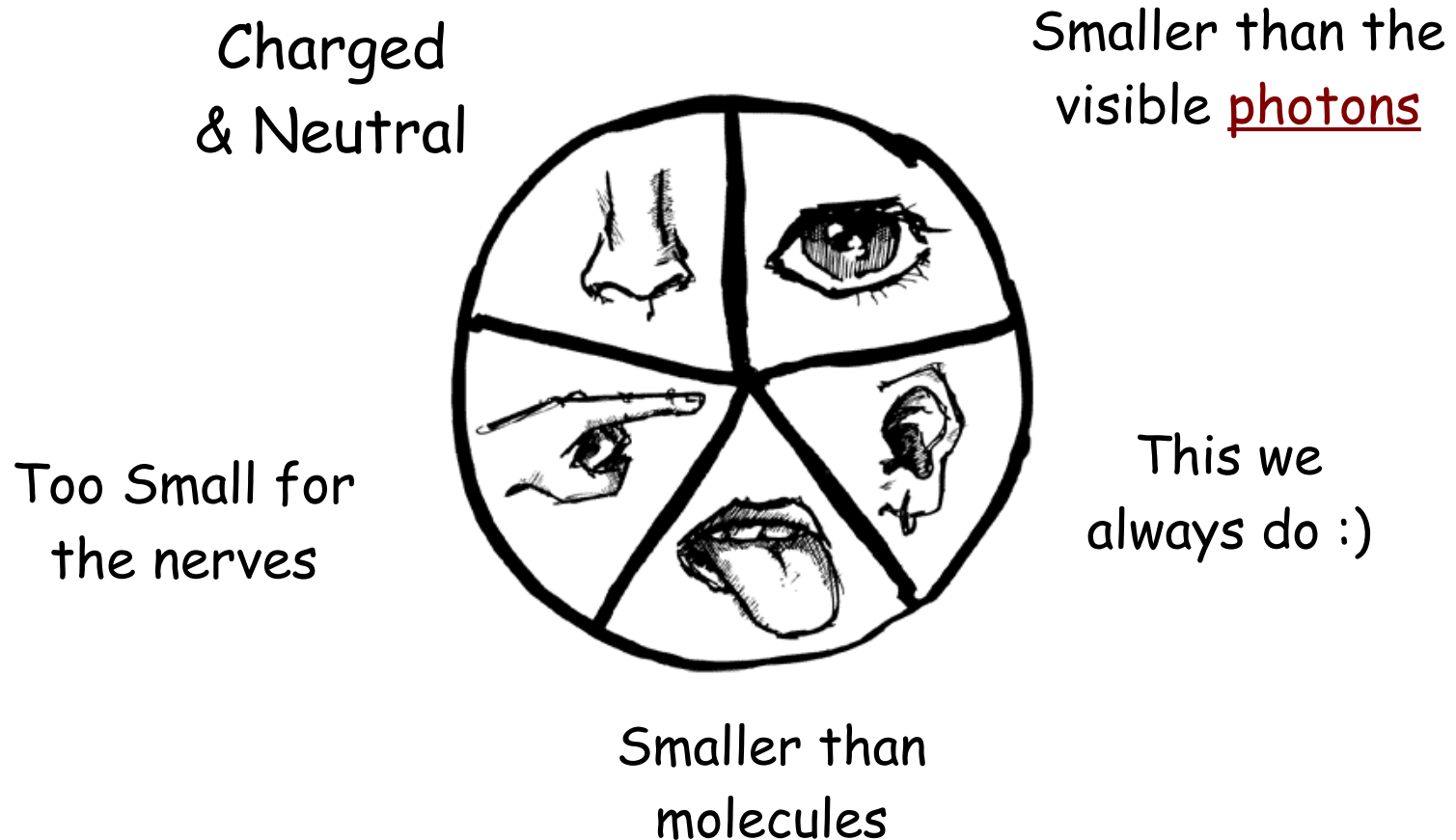
Autumn Lectures-II, 2015

Tbilisi, Georgia

- Introduction (why? where? how?)
- Working Principle of Detectors
- Scintillating detectors (organic & inorganic)
- Some Examples

Why?

Particles cannot be measured directly!



Understanding
detectors & its
artifacts

Build **new**
detector

Ph.D. Thesis,
Scientific
publications

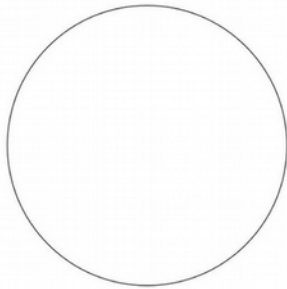


You **need** a
detector

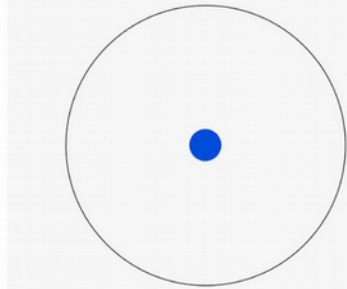
The **software**. In
fact, a part of
the detector

Keep Pushing Boundaries!

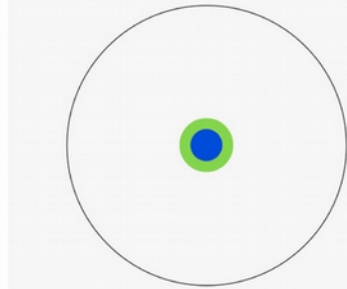
Imagine a circle that contains all of human knowledge:



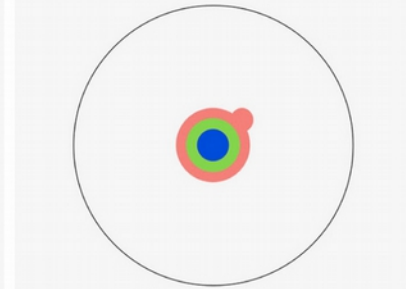
By the time you finish elementary school, you know a little:



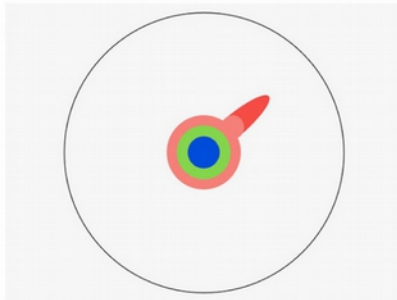
By the time you finish high school, you know a bit more:



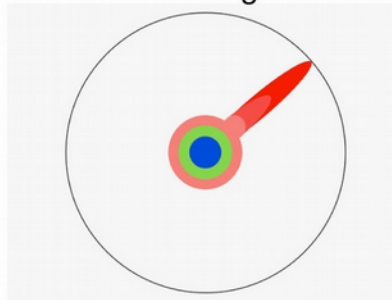
With a bachelor's degree, you gain a specialty:



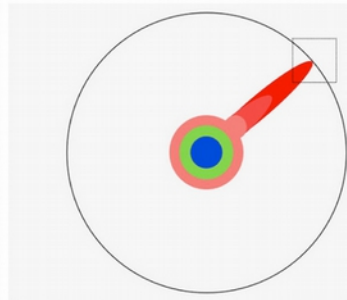
A master's degree deepens that specialty:



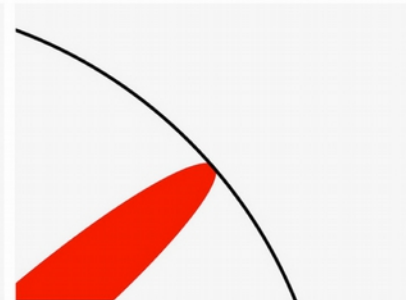
Reading research papers takes you to the edge of human knowledge:



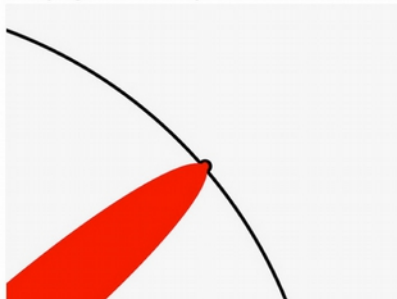
Once you're at the boundary, you focus:



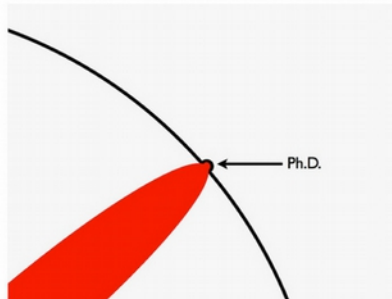
You push at the boundary for a few years:



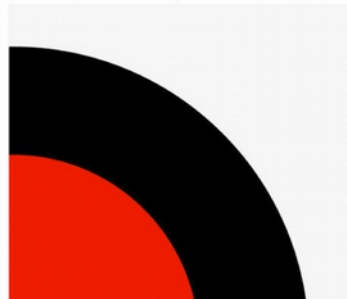
Until one day, the boundary gives way:



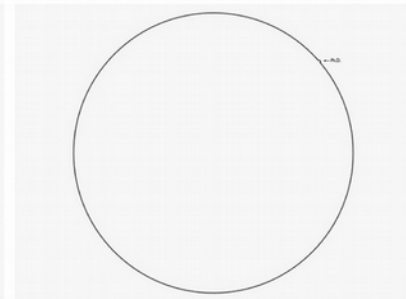
And, that dent you've made is called a Ph.D.:



Of course, the world looks different to you now:



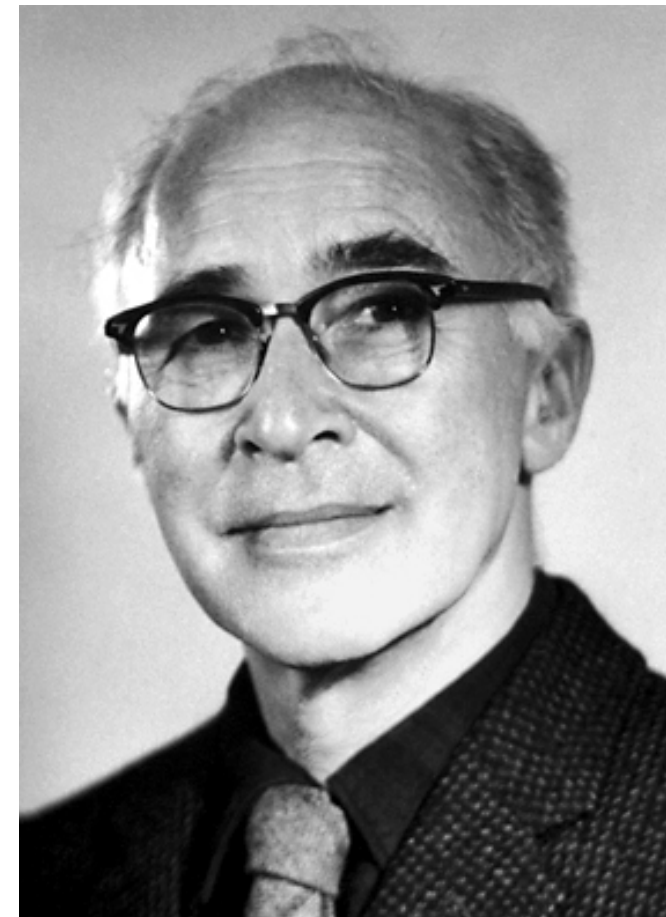
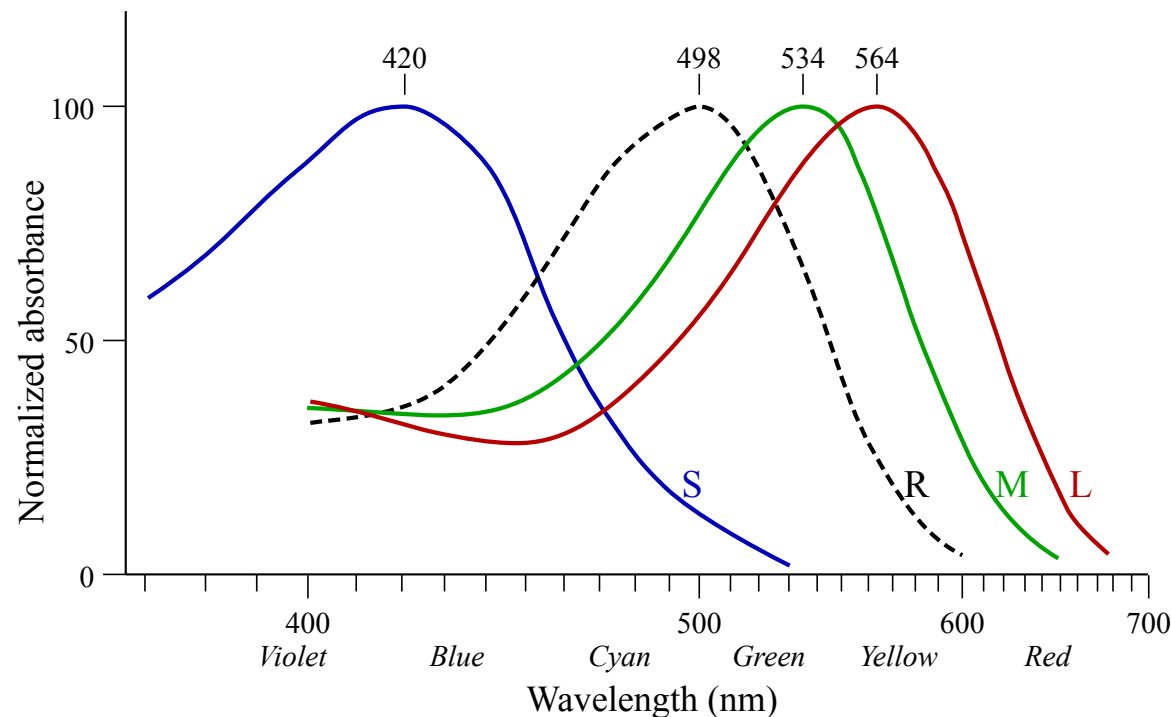
So, don't forget the bigger picture:



“A physicist is an atom's way of knowing about atoms.”

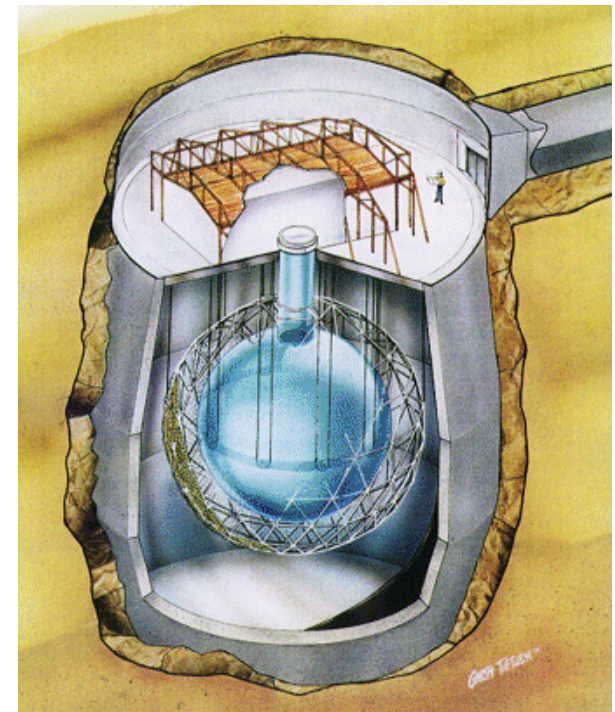
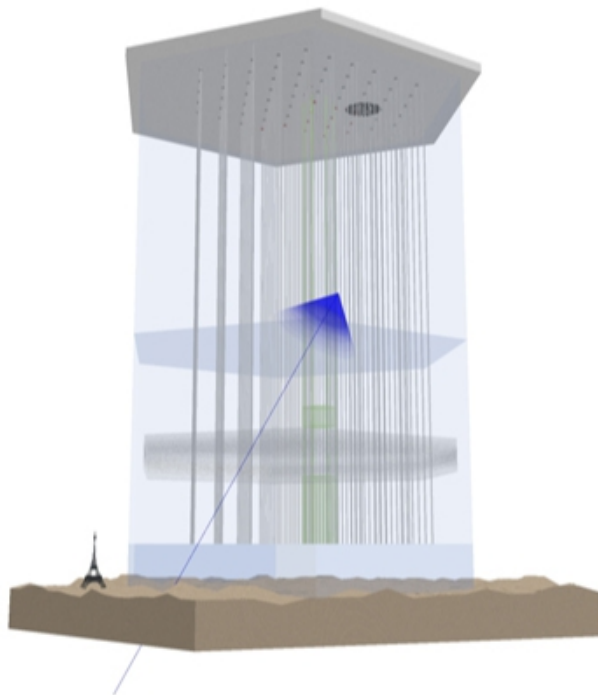
George Wald, Neurobiologist

1967 Nobel Prize



Where?

Where are Detectors on Earth?

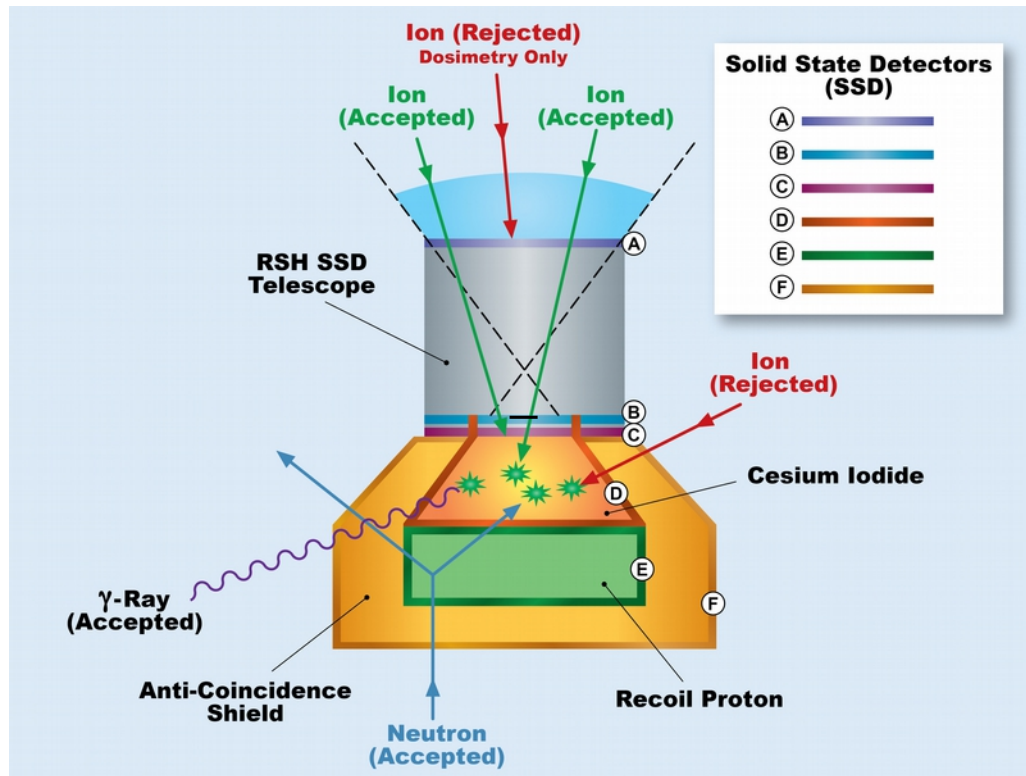


The Radiation Assessment Detector, or RAD

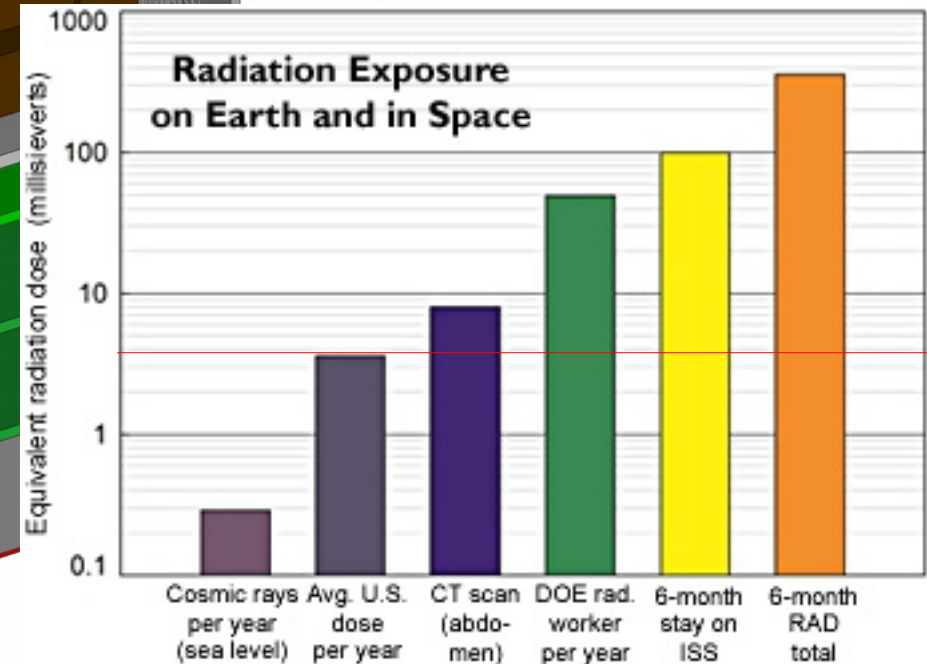


will monitor naturally occurring radiation that can be unhealthy if absorbed by living organisms. It will do so on the surface of Mars, where there has never before been such an instrument, as well as during the trip between Mars and Earth.

Curiosity's Radiation Assessment Detector



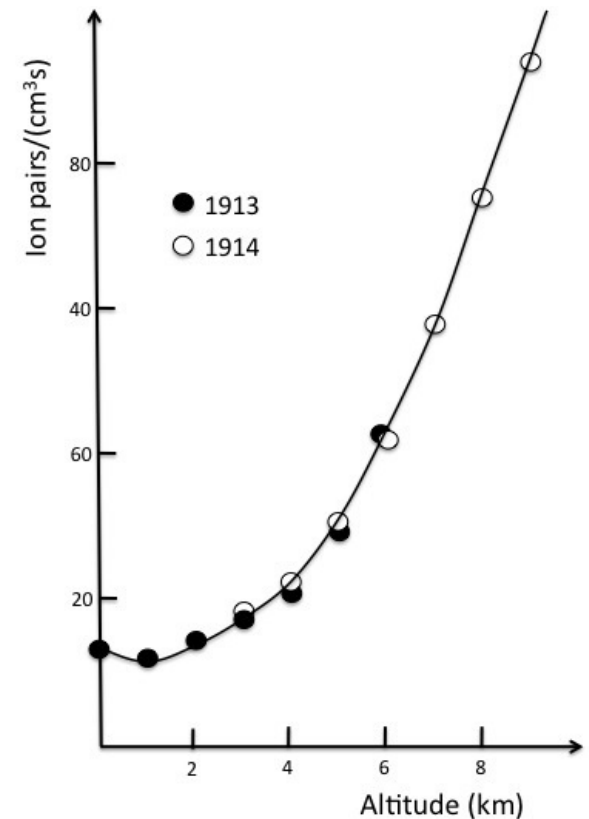
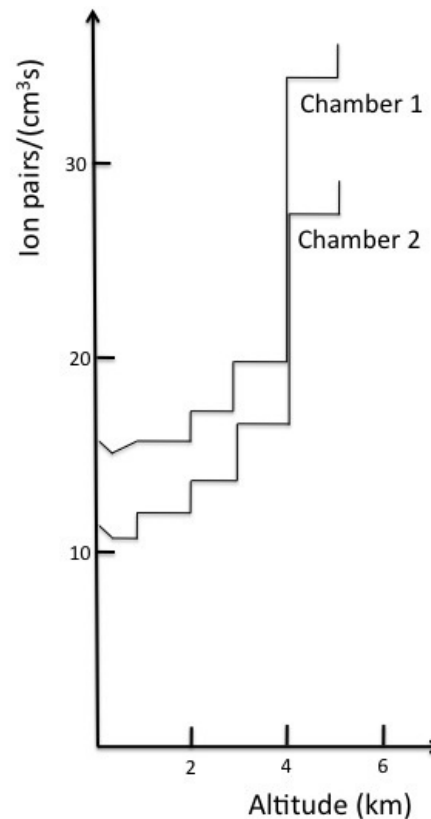
NASA scientists reported that a possible manned mission to Mars may involve a greater radiation risk than previously believed

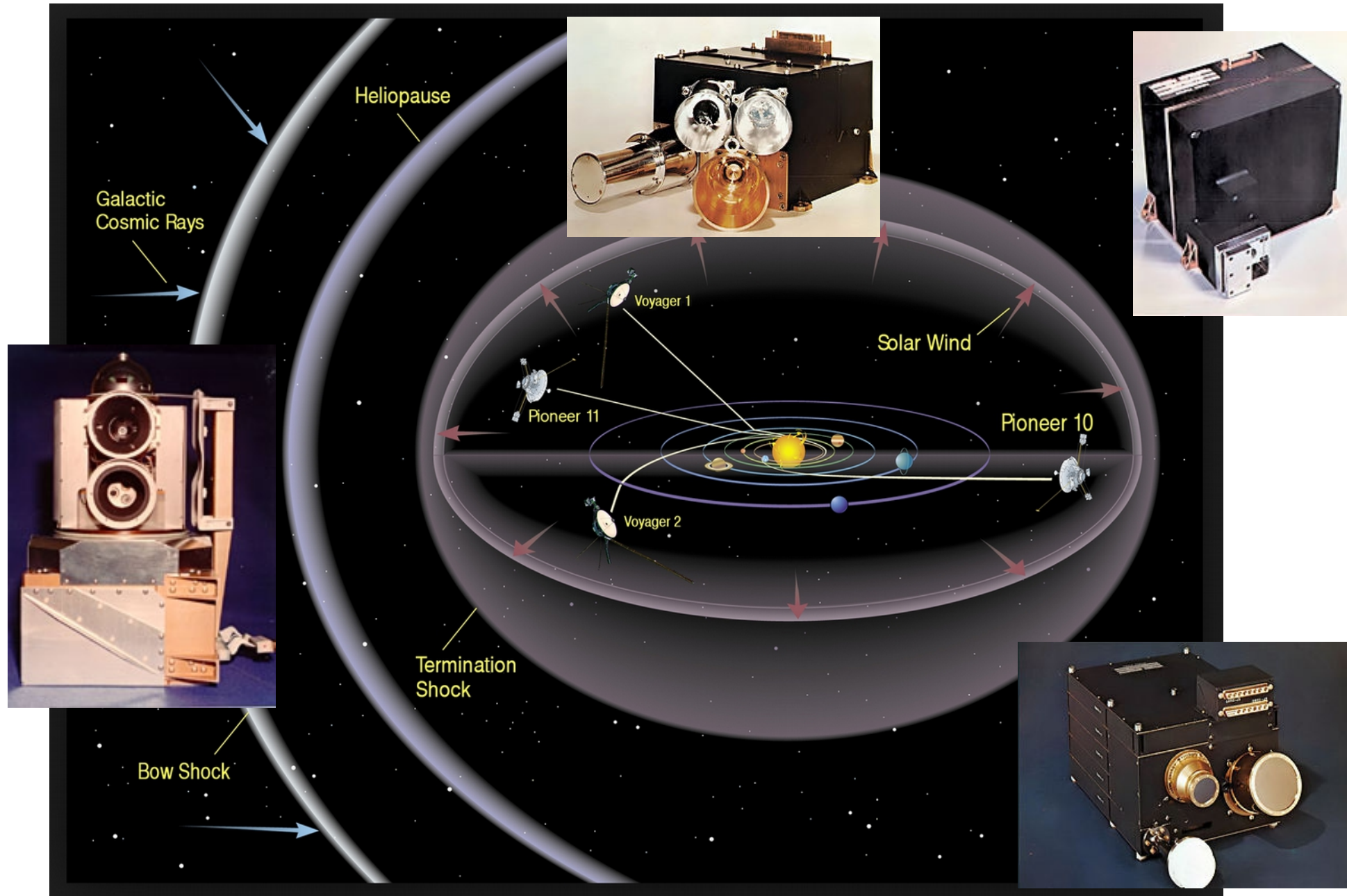




Victor Hess received the
Nobel Prize in Physics in 1936 for
his discovery

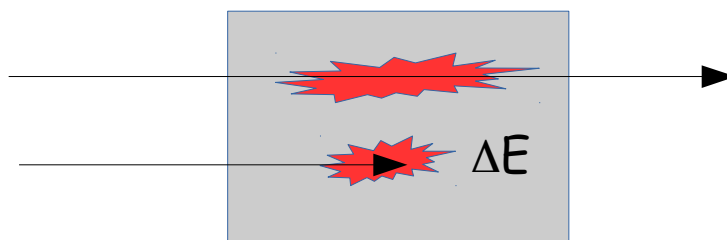
The Hess balloon flight took place on 7 August 1912. By sheer coincidence, **exactly 100 years later** on 7 August 2012, the Mars Science Laboratory rover used its RAD instrument to begin measuring the radiation levels on another planet for the first time.





How & What?

Only the result of an interaction with matter
will be observed



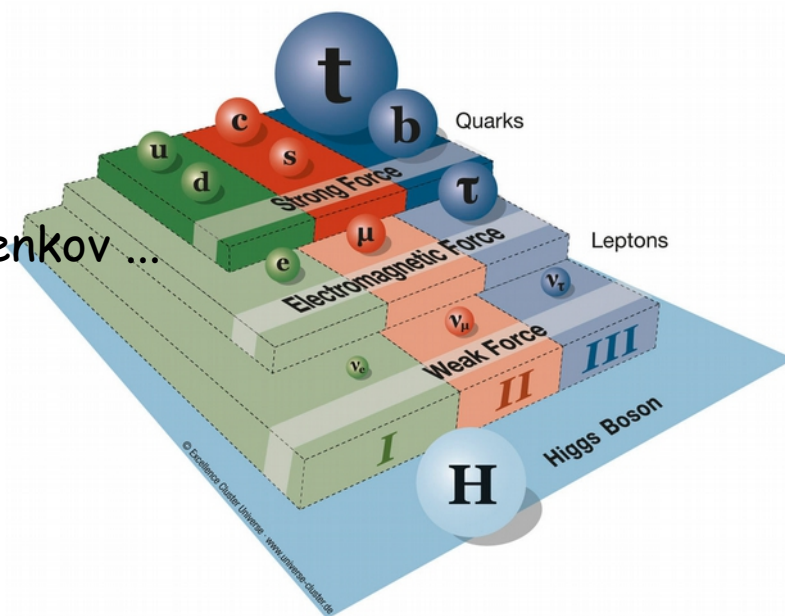
The detection of particles happens

via their ΔE in the material it traverses ...

Charged particles: Ionization, Bremsstrahlung, Cherenkov ...

Photons: EM $\rightarrow$ Photo, Compton, pair production

Hadrons: Strong, EM, Weak interactions



HEAVY CHARGED PARTICLES - STOPPING POWER II

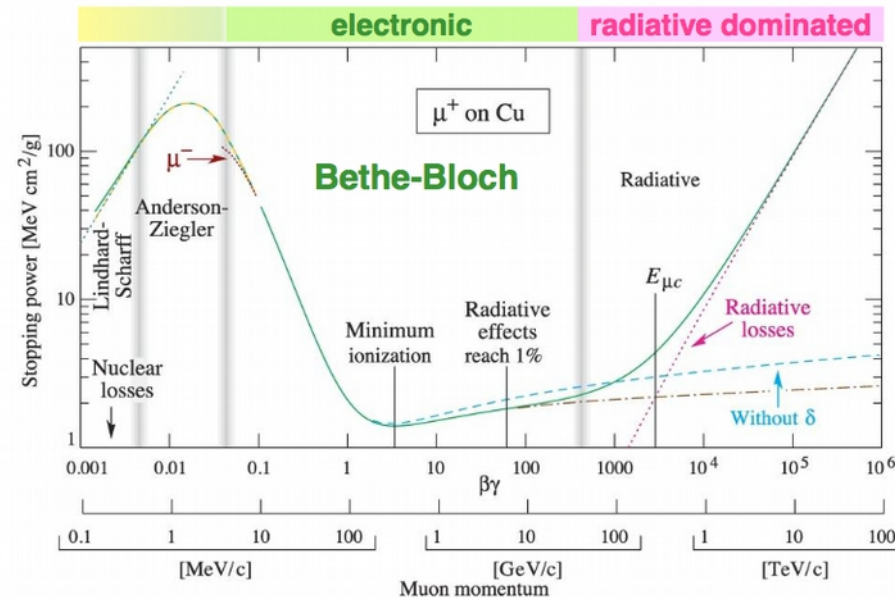
Bethe-Bloch range

$\alpha < \beta < 0.1$: β_{electron} neglected

stopping power

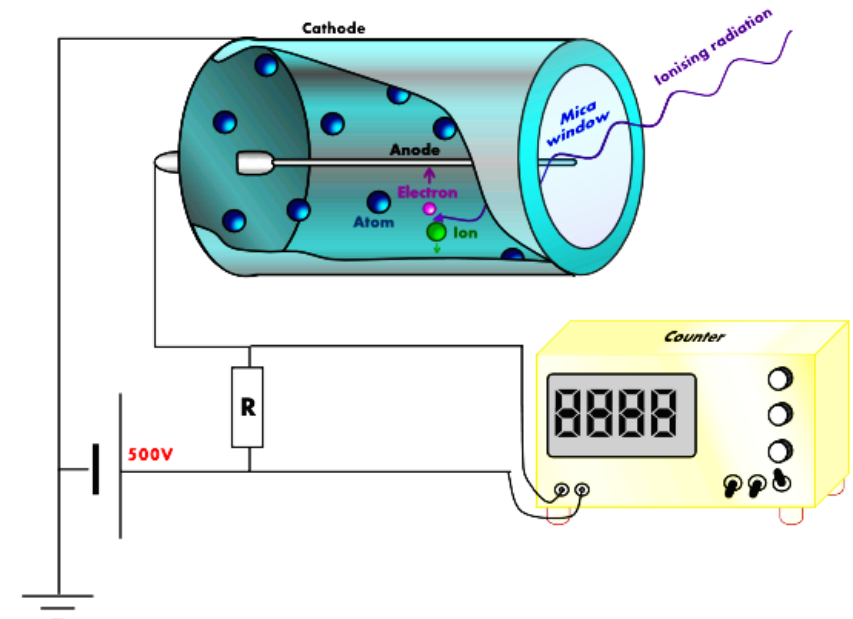
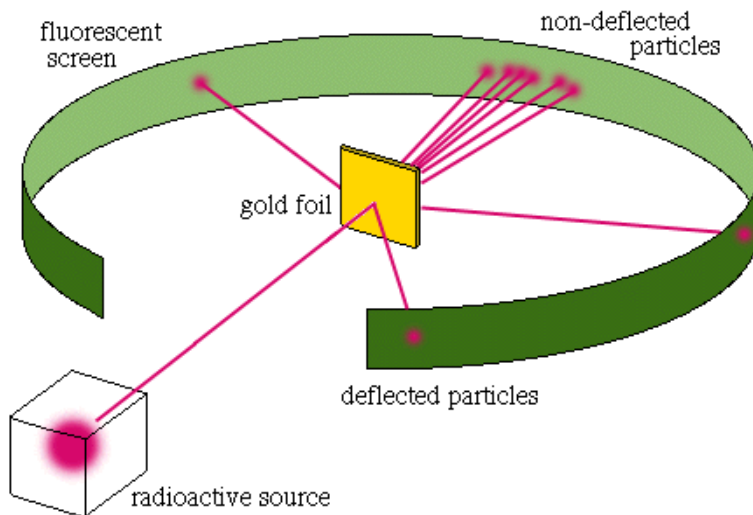
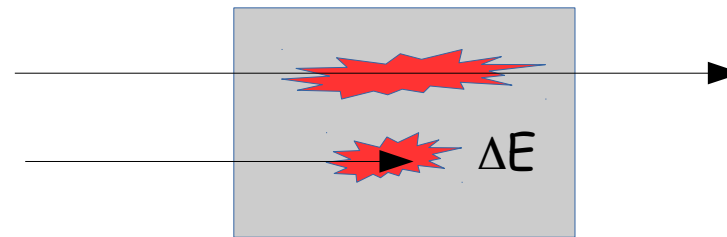
$$S = \left(-\frac{dE}{dx} \right) \cdot \frac{1}{\rho}$$

[MeV · cm²/g]

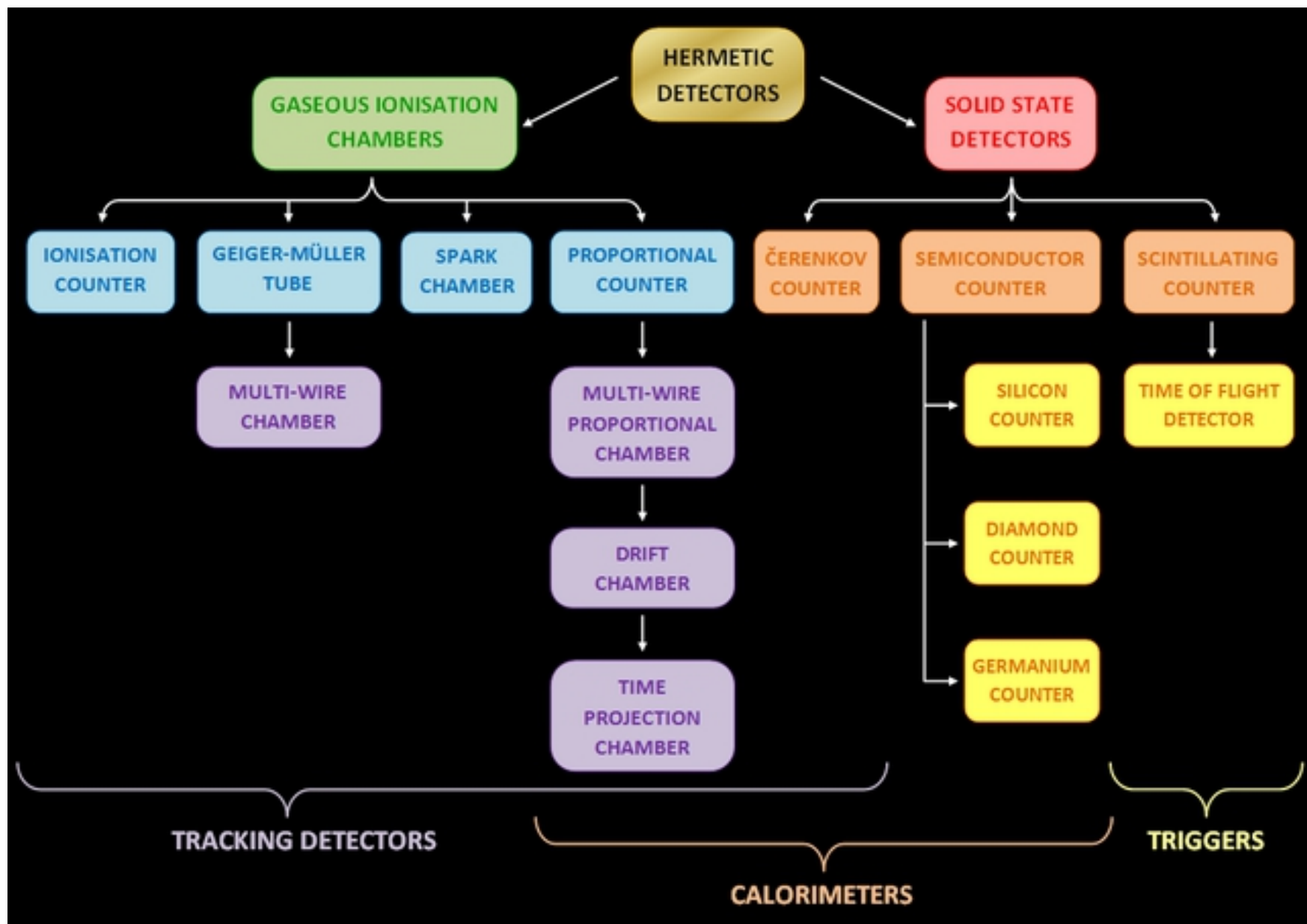


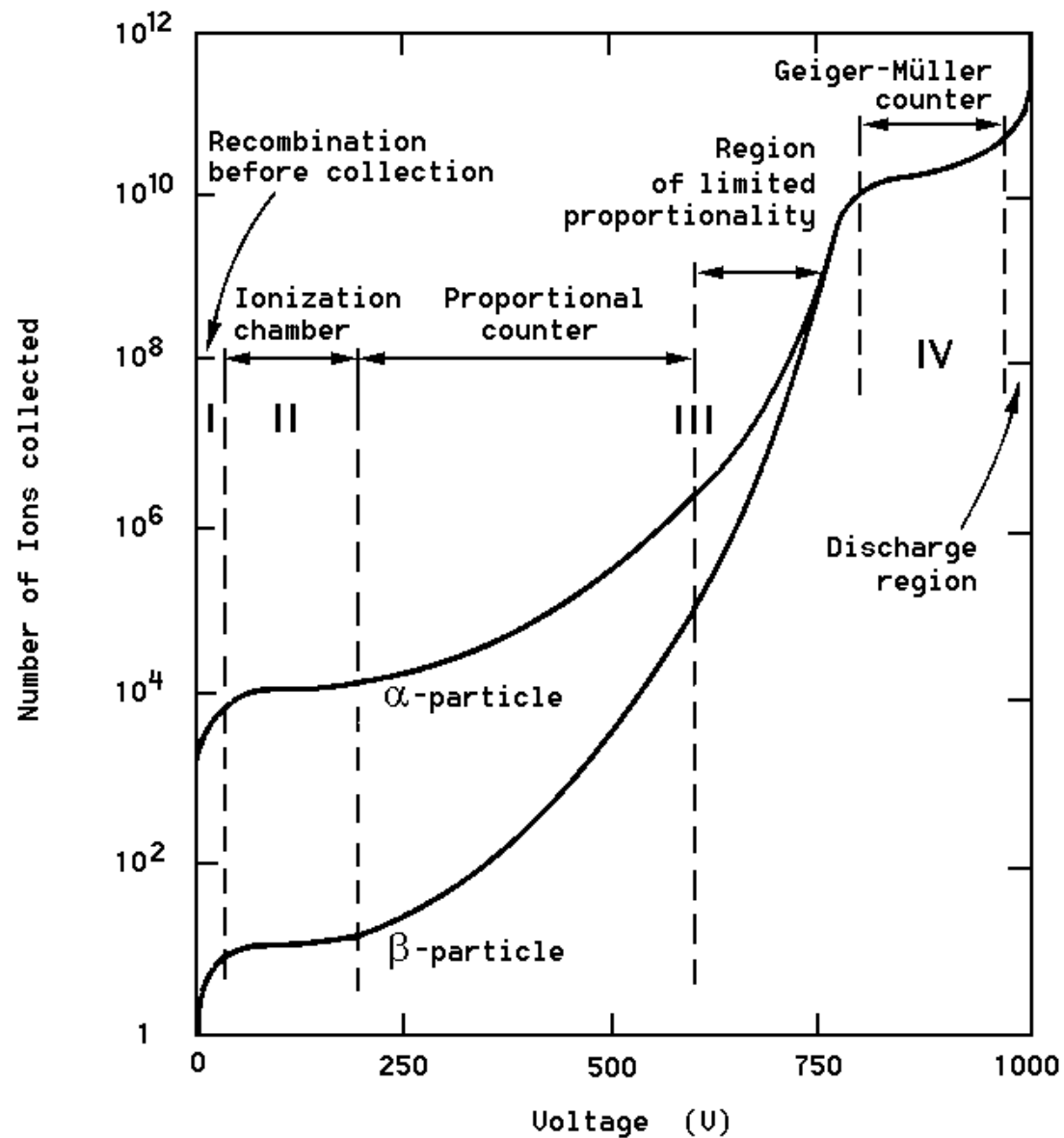
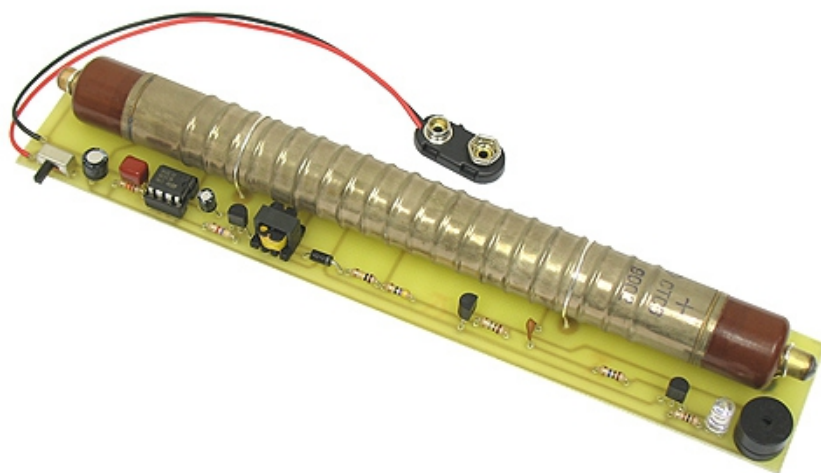
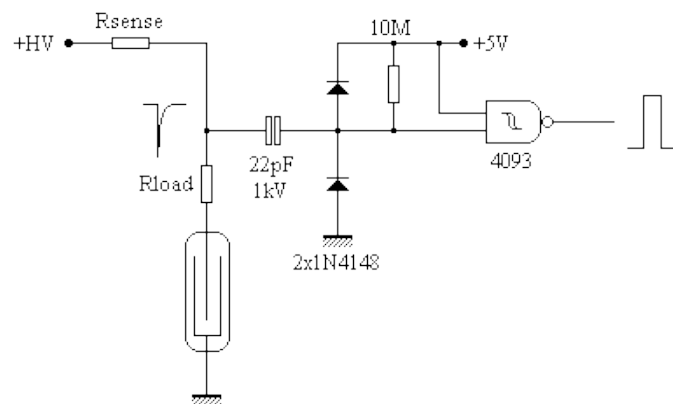
*C₀ shell corrections
for low energies*

$$S = 4\pi N_A r_e^2 m_e c^2 \frac{Z_{\text{target}}}{A} \cdot \frac{z^2_{\text{projectile}}}{\beta^2} \cdot \left[\frac{1}{2} \ln \left(\frac{2m_e c^2 \beta^2 \gamma^2 T_e^{\text{max}}}{I^2} \right) - \beta^2 - \frac{C_0(\beta^2)}{Z_{\text{target}}} - \frac{\delta(\beta\gamma)}{2} \right]$$

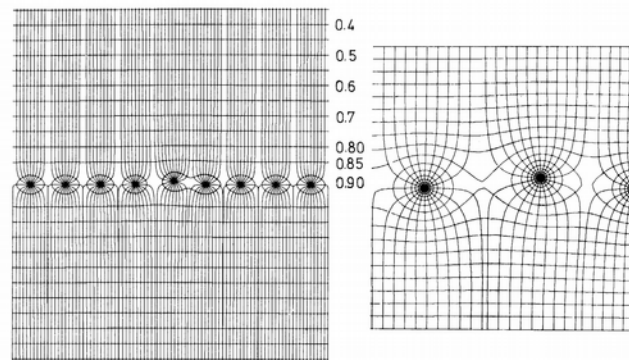
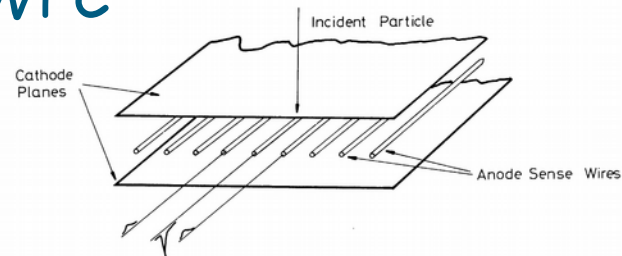
[illegible]

- Ionizing (Gaseous & Semiconductor)
- Scintillating (Organic & Inorganic)
- Cherenkov (VETO, RICH & DIRC)

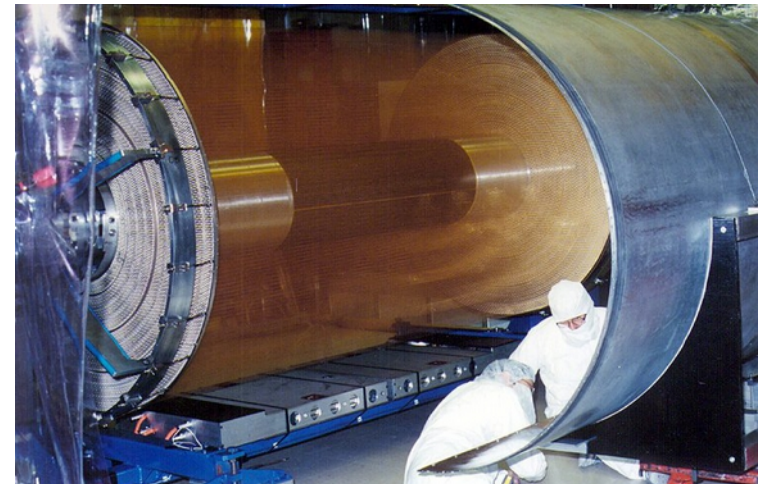
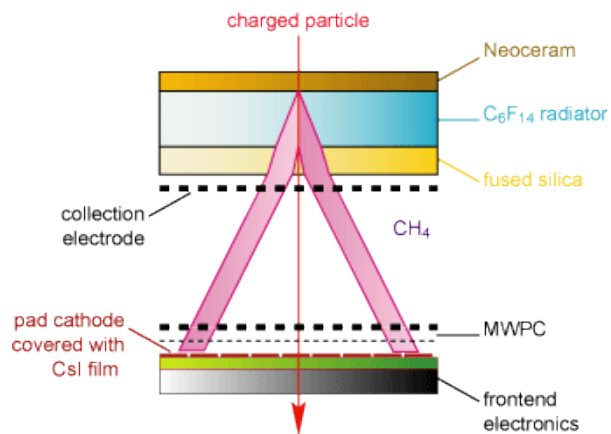
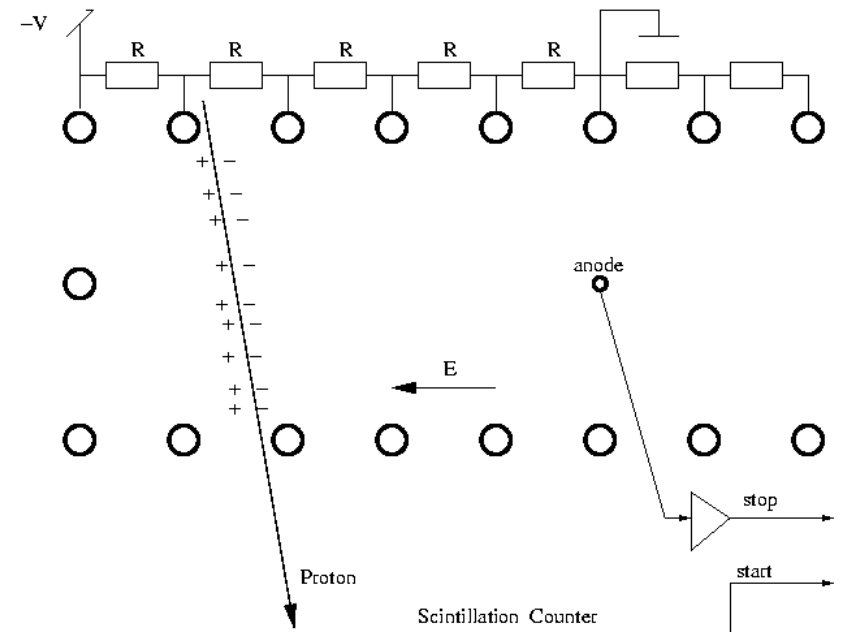


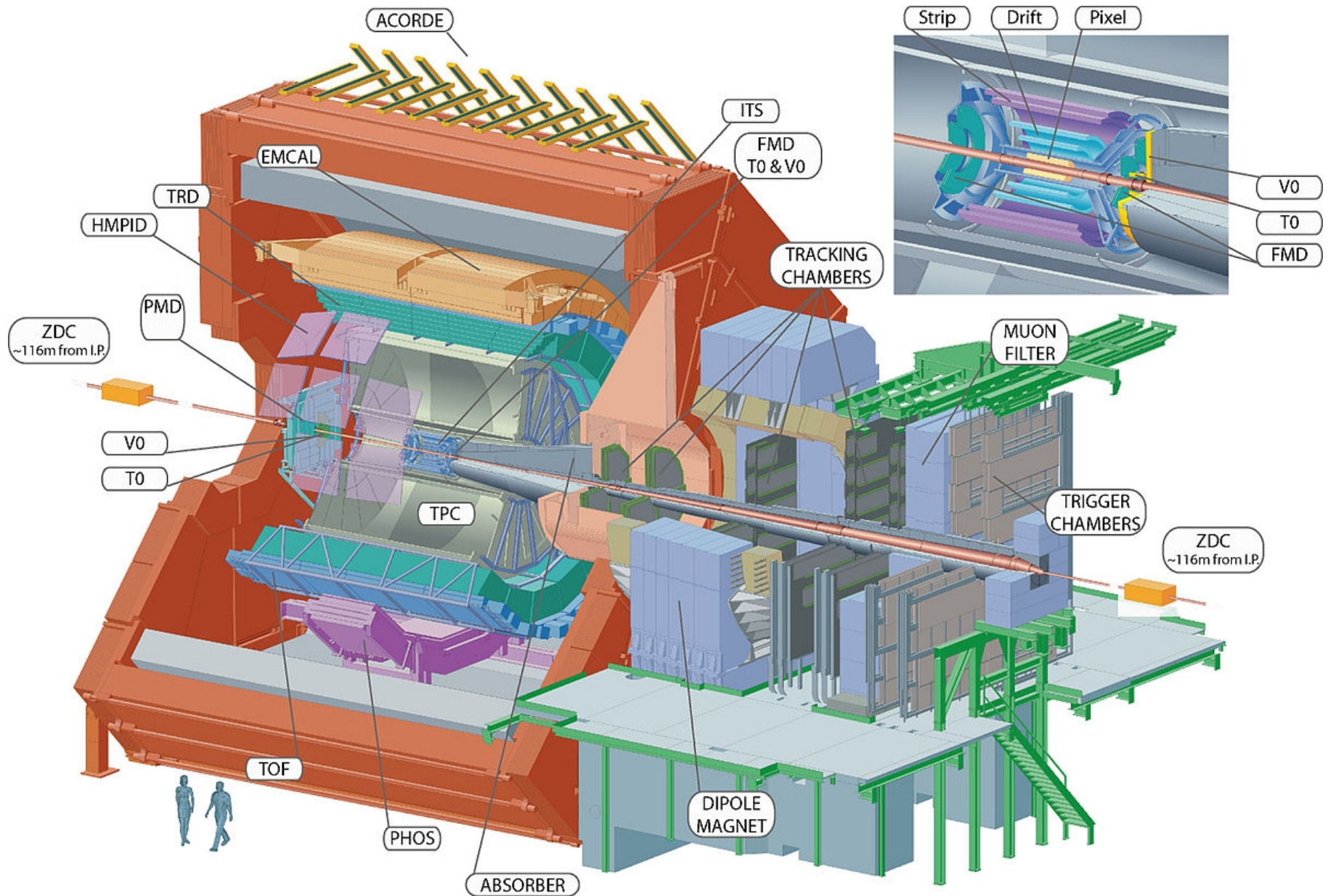


MWPC



MDC

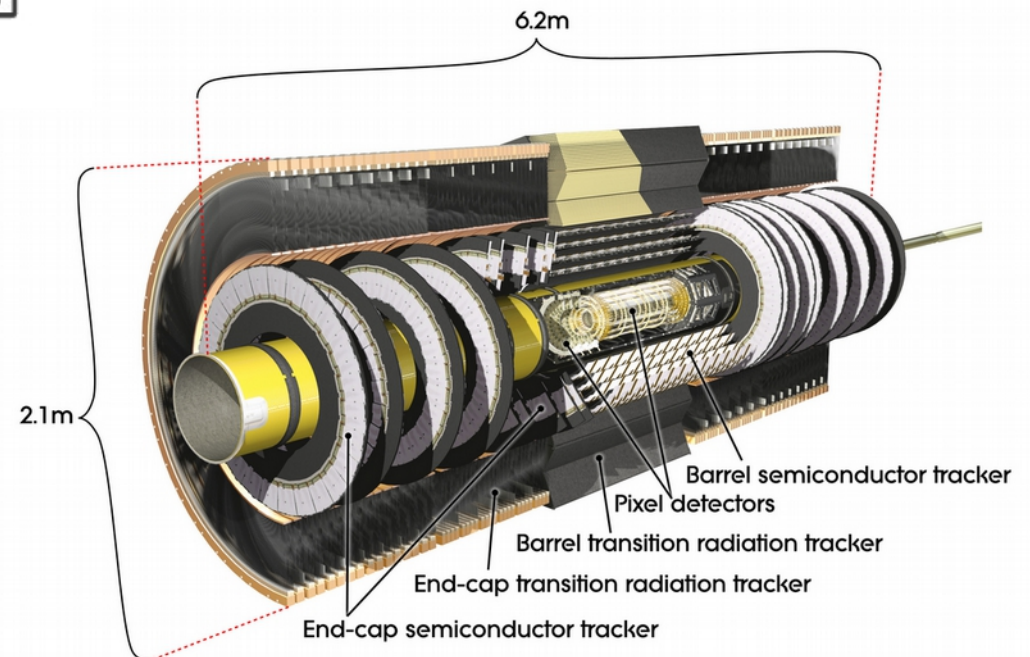
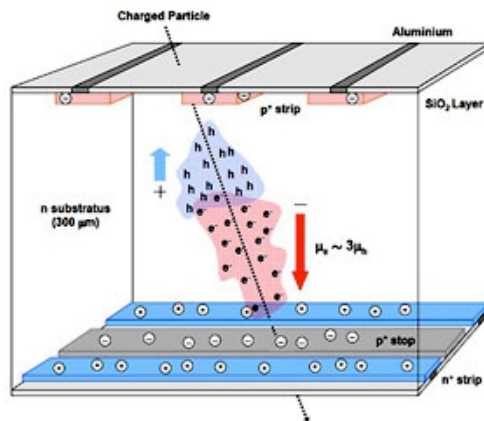
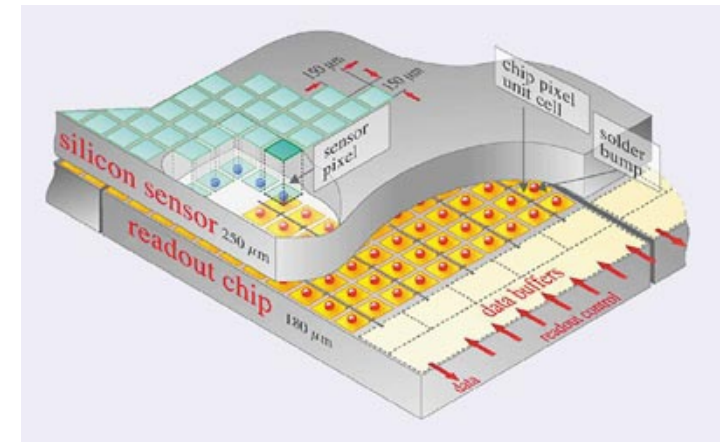
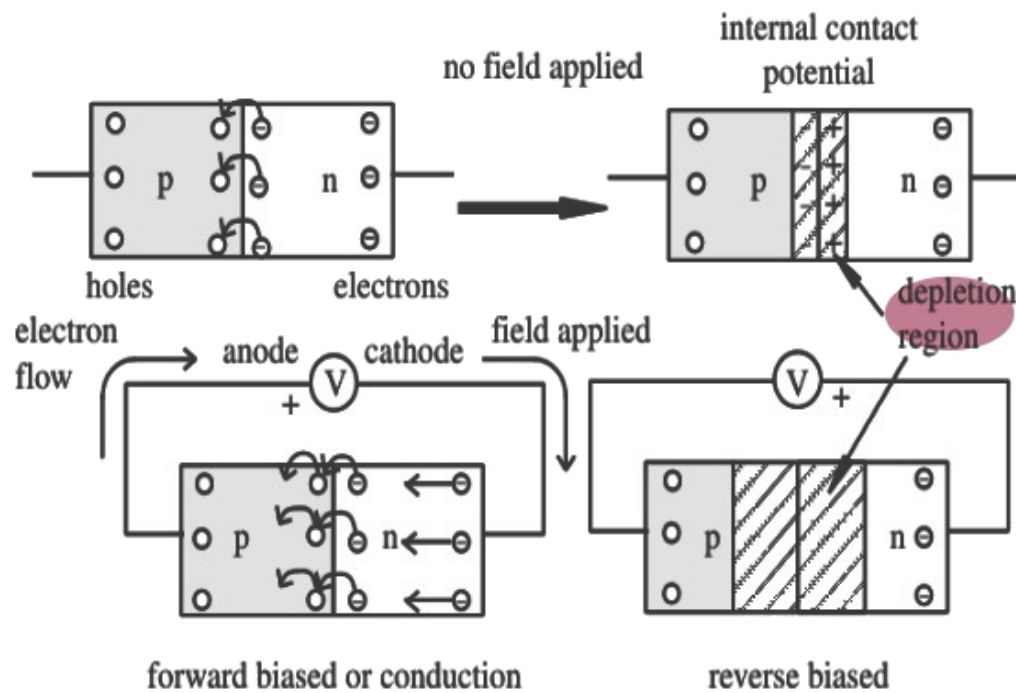




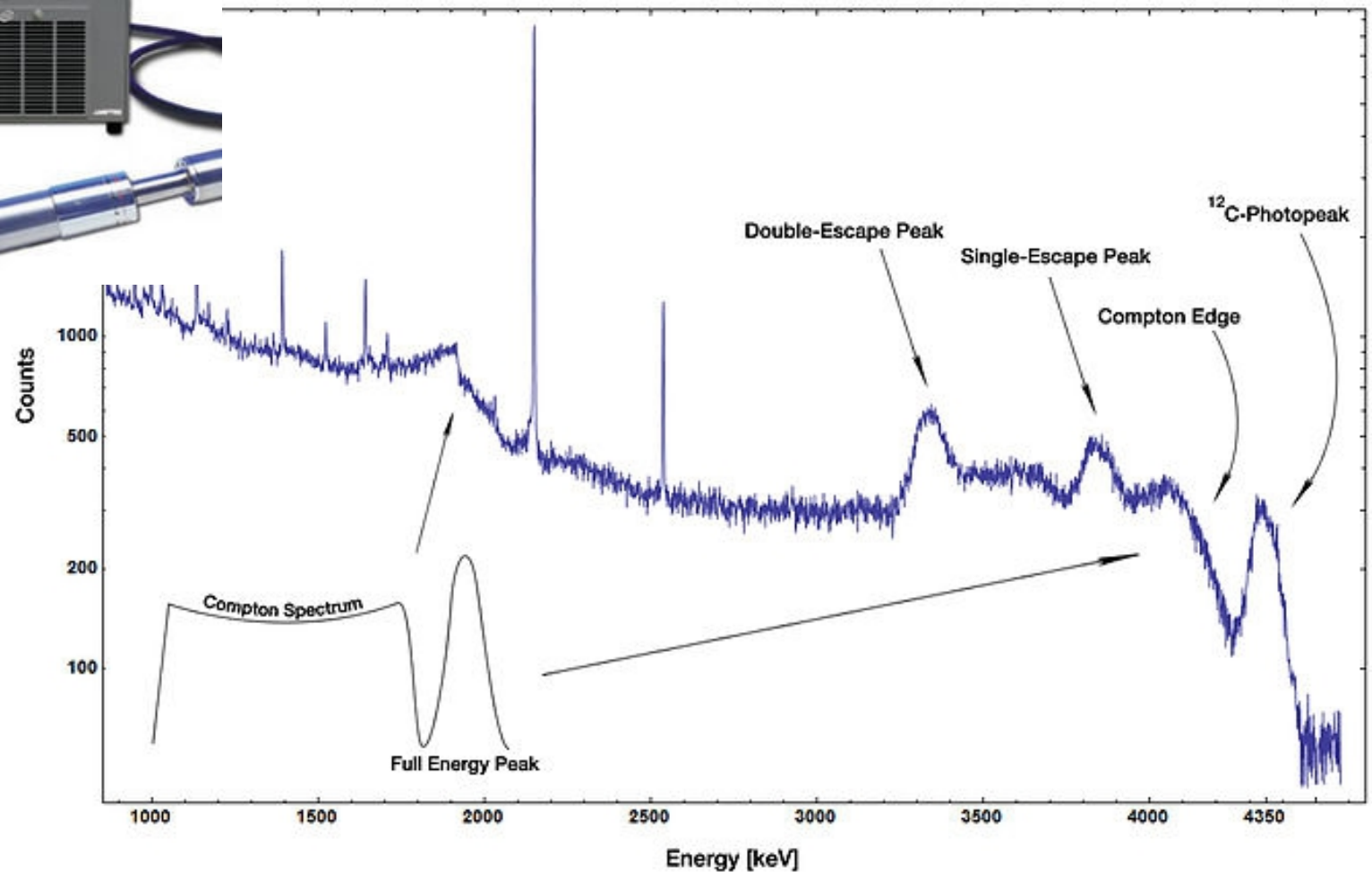
Semiconductors



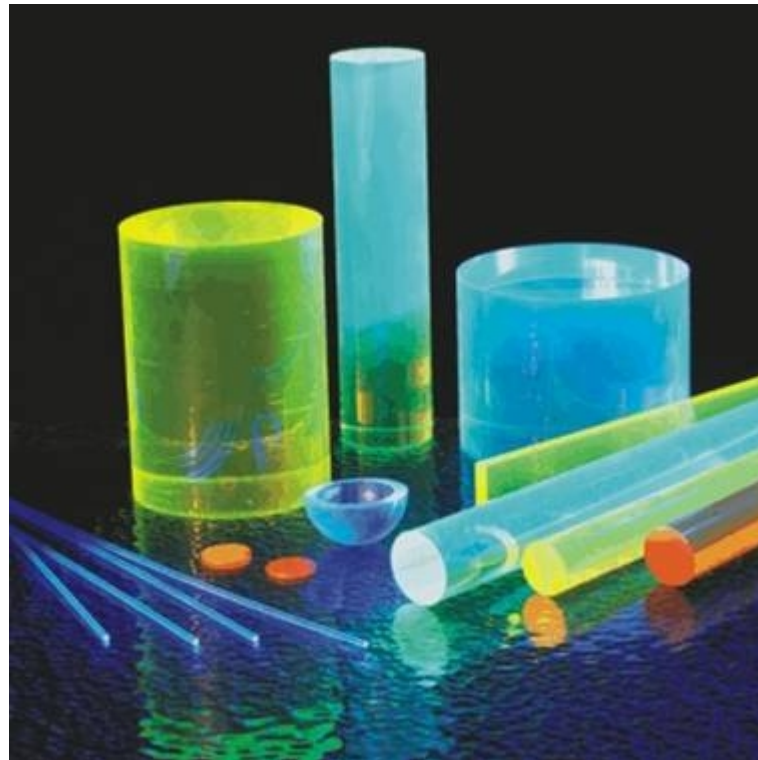
Working Principle

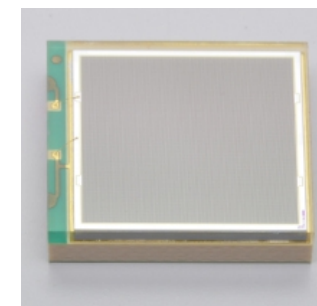
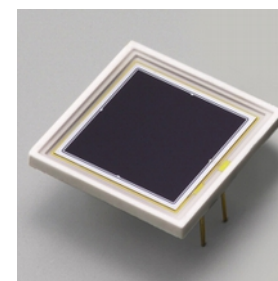
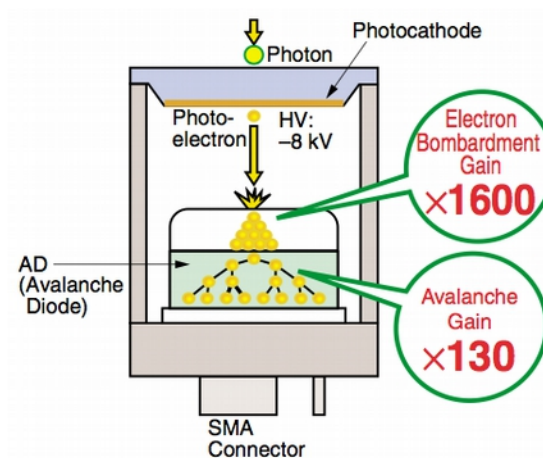
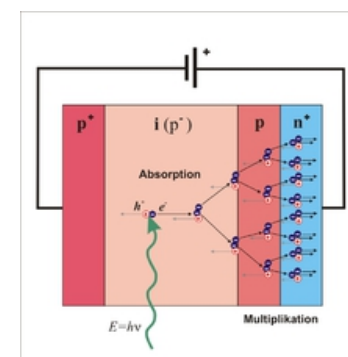
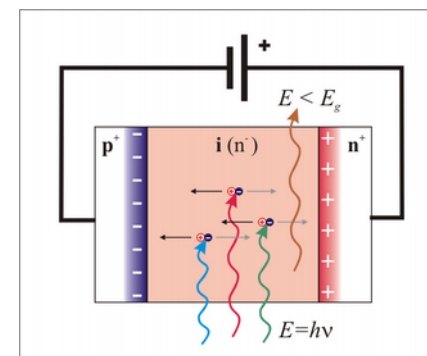
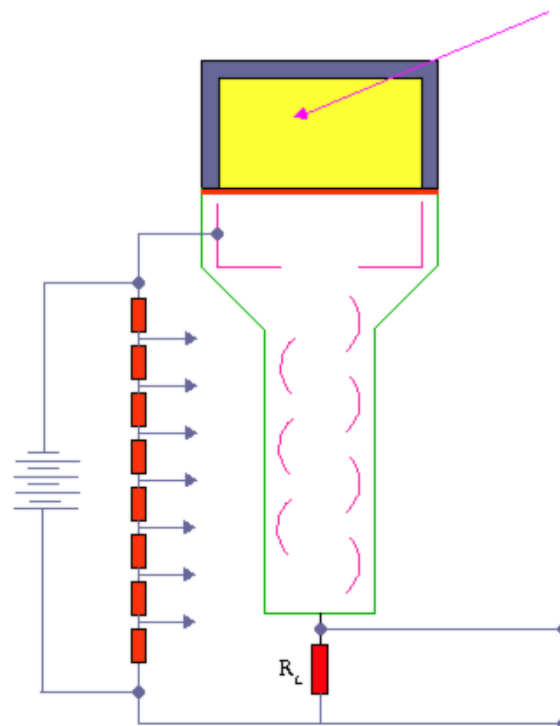
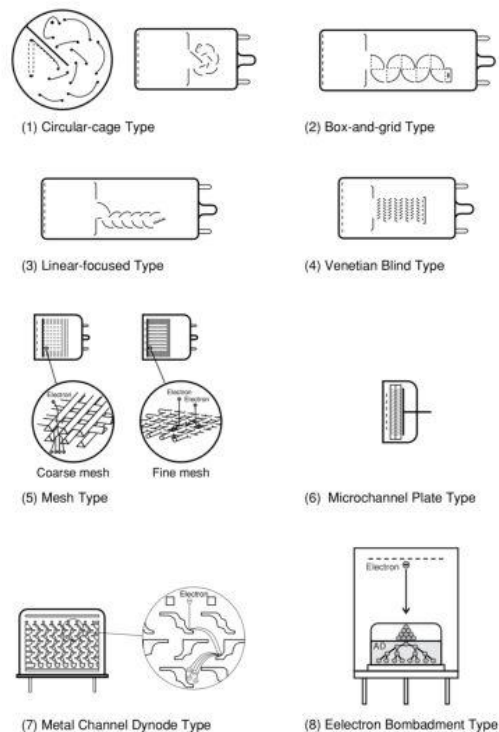


How to Interpret the Spectrum

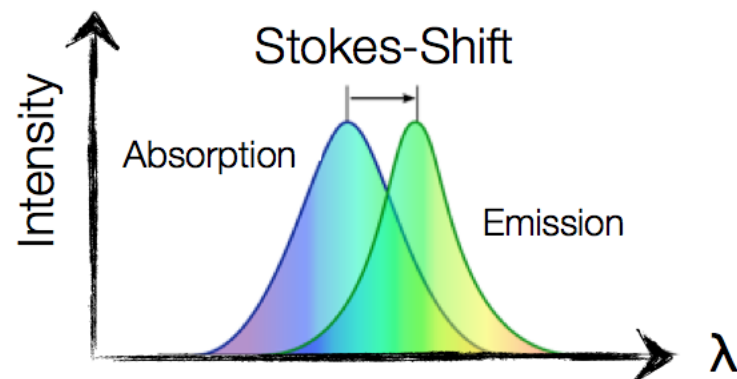
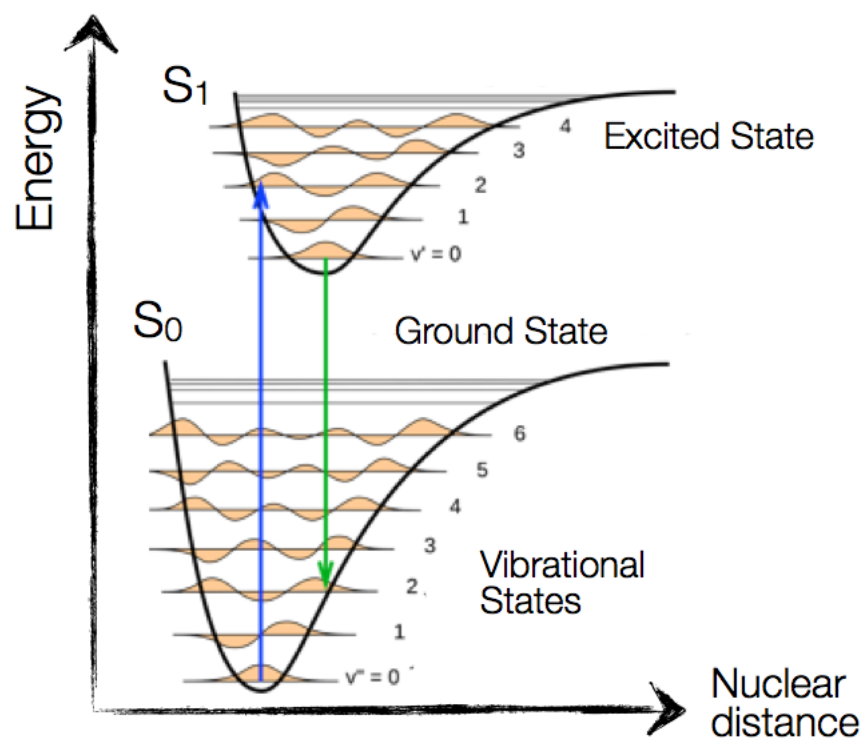


Scintillators





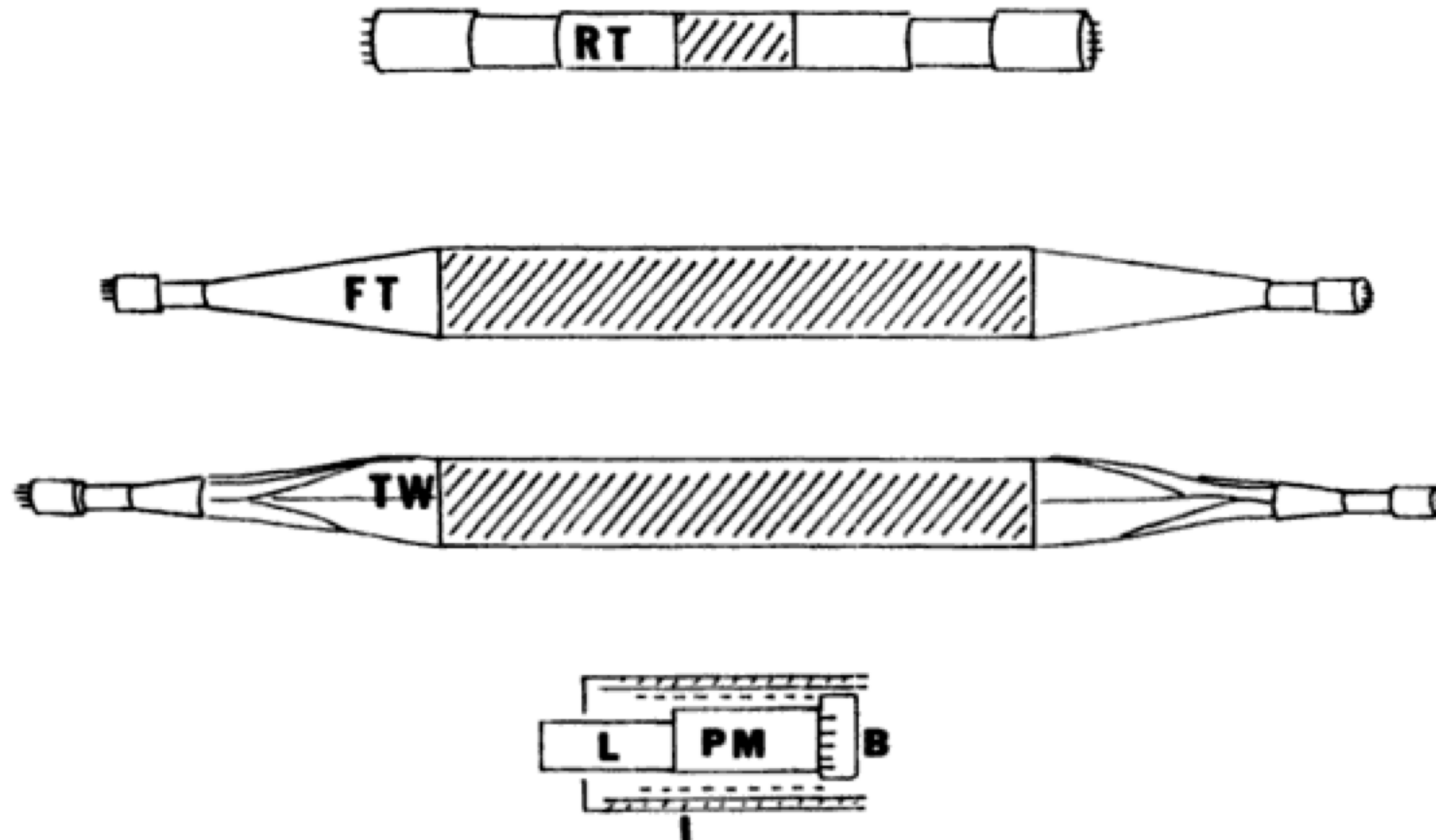
The light emission is governed by the electronic transitions in the molecule. The electronic levels have a typical energy spacing of ~ 4 eV. The vibrational levels of the molecule ($dE \sim 0.2$ eV) also play a role. Electrons in high levels typically deexcite to the lowest excited state without emission of radiation.



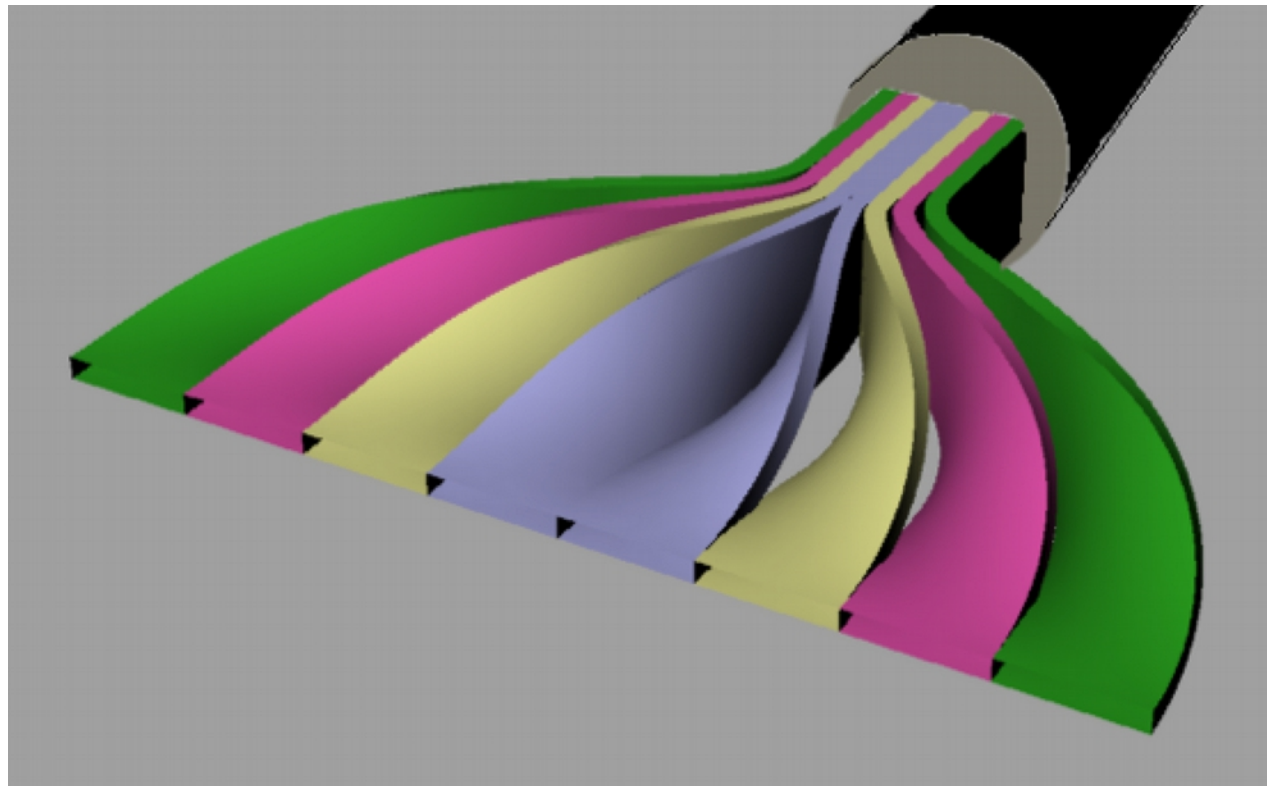
Types: (RT) rectangular, (FT) "fish tail," (TW) two twisted strips,
The bottom figure shows a photomultiplier

tube assembly (PM) with base (B), standard coupling piece of light guide (L), mu metal shield (dotted lines), and soft iron shield (I).

(After G. D'Agostini et al., Nuc. Instr. Meth, 185: 49, 1981.)

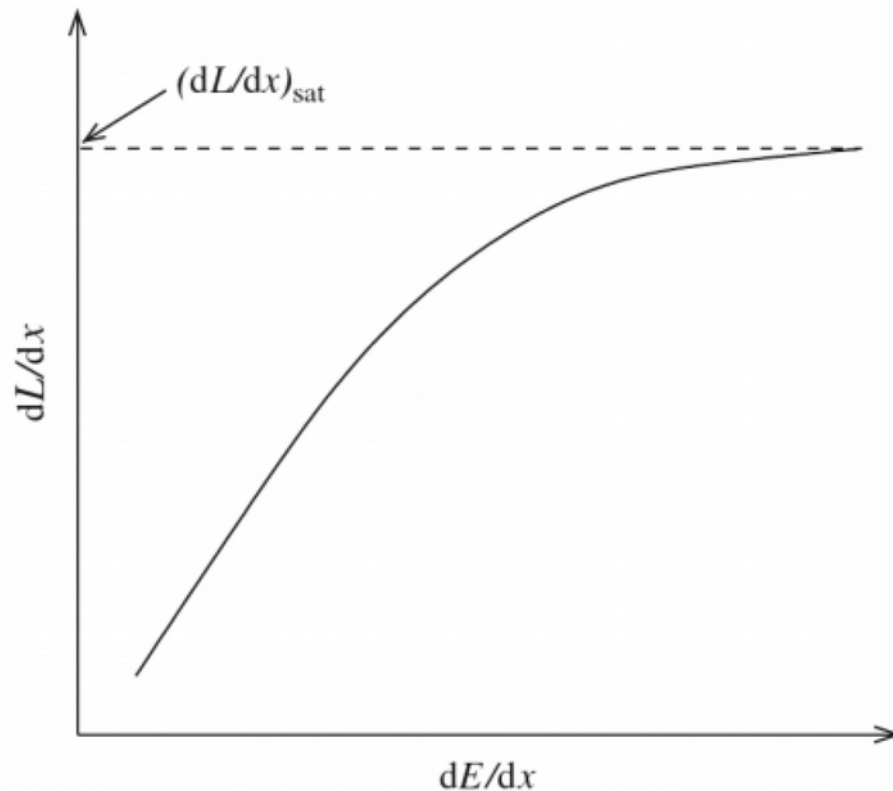


- presents a fundamental limitation on the transmission of light through the pipe. The theorem requires that the light flux per unit solid angle not increase while propagating in the pipe. This has the consequence that the maximum fraction of light that can be transmitted is A_{pm}/A_{sc} , where A_{pm} (A_{sc}) is the area of the PMT face (scintillator edge) and $A_{pm} < A_{sc}$.



The amount of light produced for a given energy loss is not constant but depends on the production of quenching centers. These are activated molecules raised to excited vibrational levels. The number of produced photons roughly follows the equation

$$dL/dX = \frac{A \cdot dE/dX}{1 + B \cdot dE/dX}$$



where A and B are constants and dE/dx is the ionization energy loss. The parameter B is sometimes referred to as the a to b ratio. **Note that the light output saturates for large energy losses.**

The fraction of the emitted light transmitted via total internal reflection in one direction along the slab is

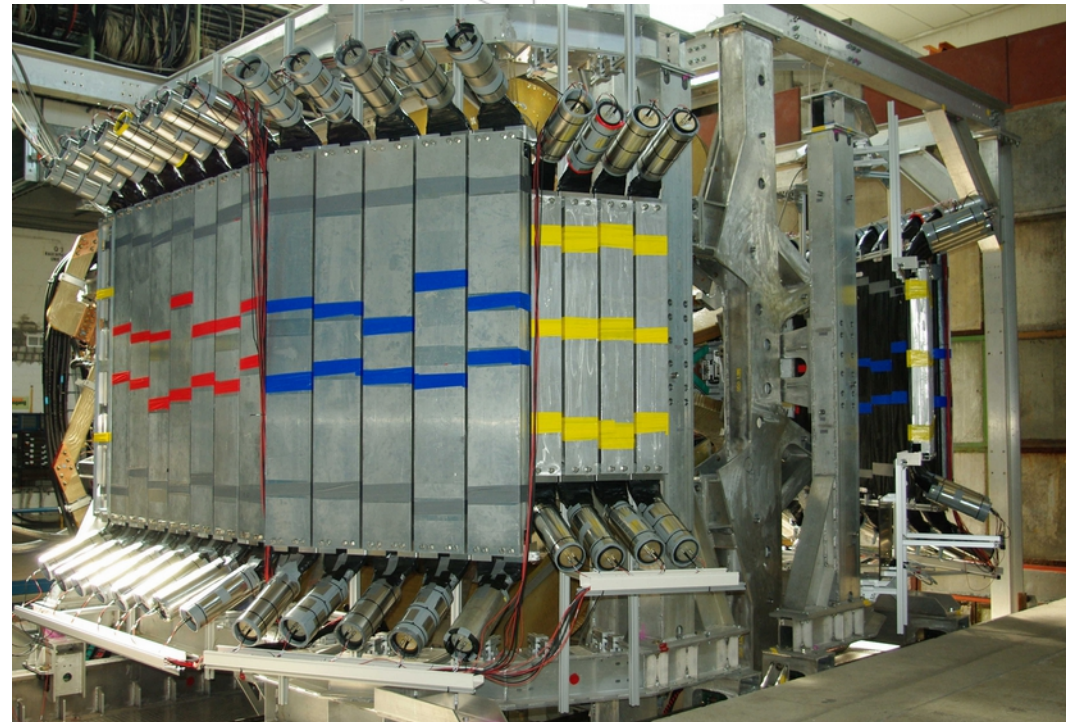
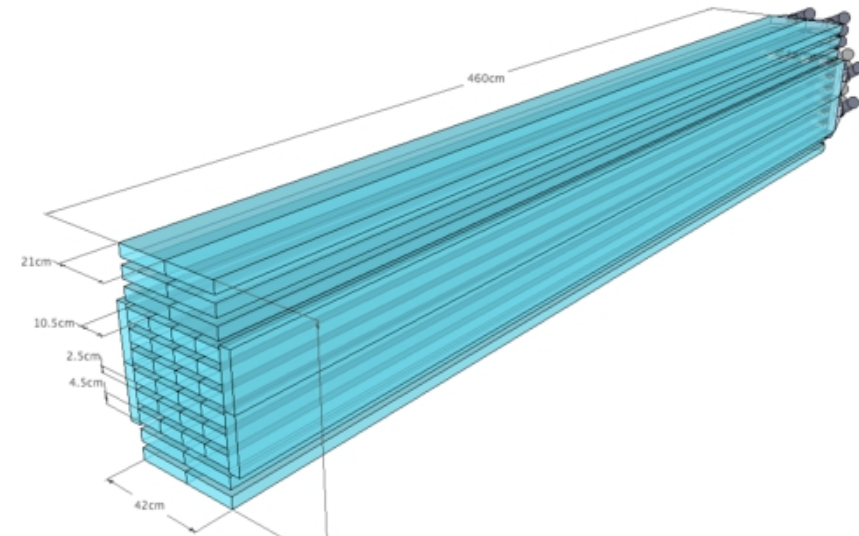
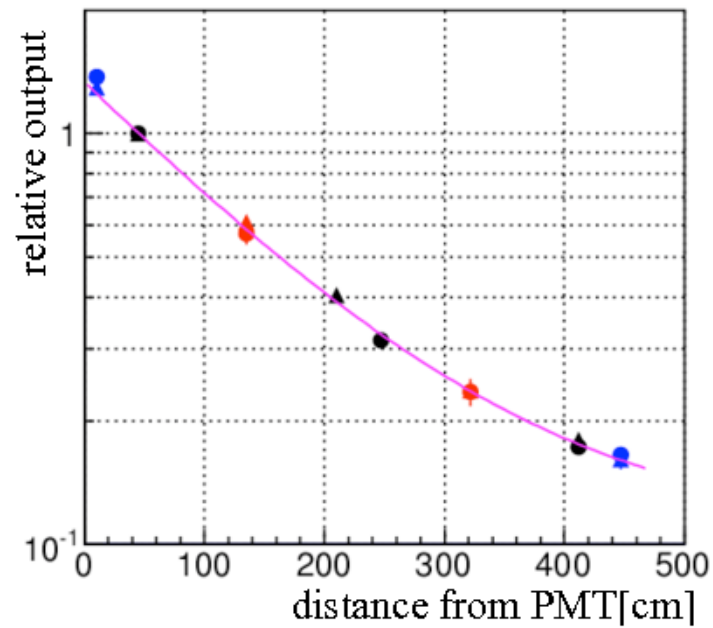
$$f = \frac{1}{4\pi} \int_0^\theta 2\pi \cdot \sin \theta d\theta = \frac{1}{2} \left(1 - \frac{1}{n} \sqrt{n^2 - 1} \right)$$

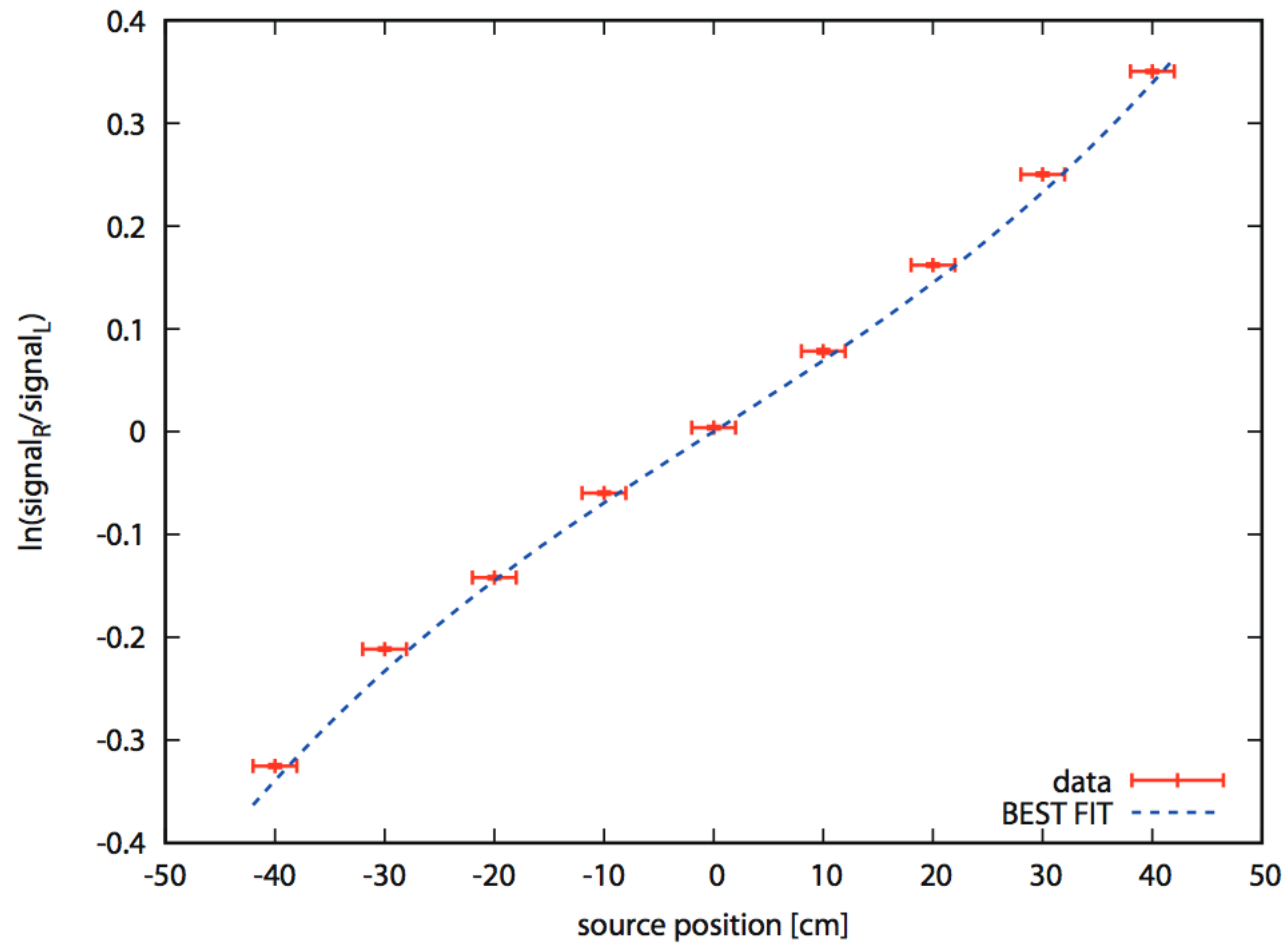
Note that this result is independent of the slab dimensions. For typical plastic scintillator with $n = 1.58$, Eq. gives $f = 0.113$.

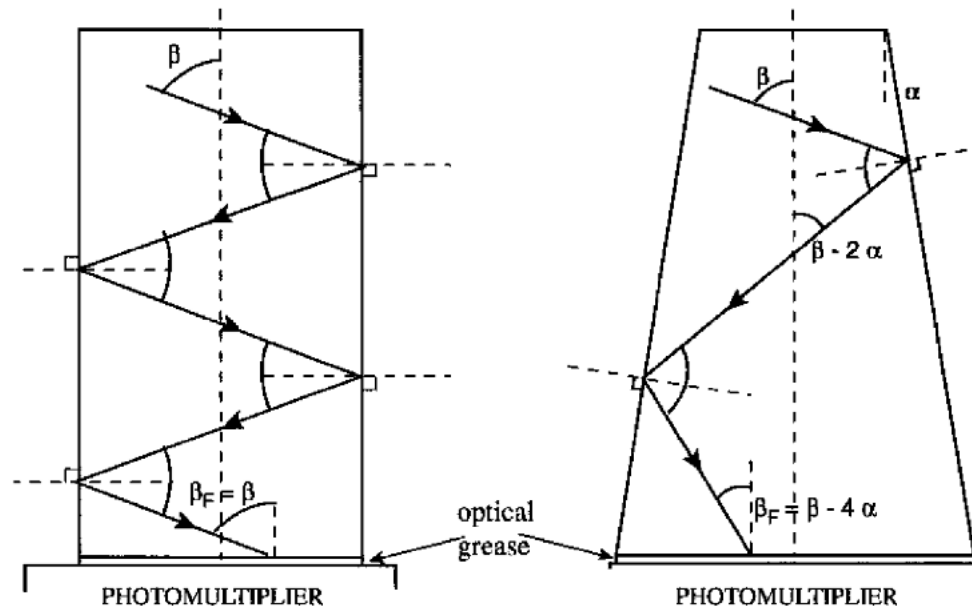
An improvement in light transmission in small counters can be achieved by using a reflector on the end of the scintillator opposite to the collection direction.

The reflectors may be either specular or diffuse and may be in direct contact with the medium or separated by a small air gap. Reflectors on the other four sides of the slab will in general have a much smaller influence on the collection efficiency.

$$Y = Y_0 \cdot e^{\frac{-l}{\lambda}}$$







Rough design values

Energy loss in plastic = 2 MeV/cm

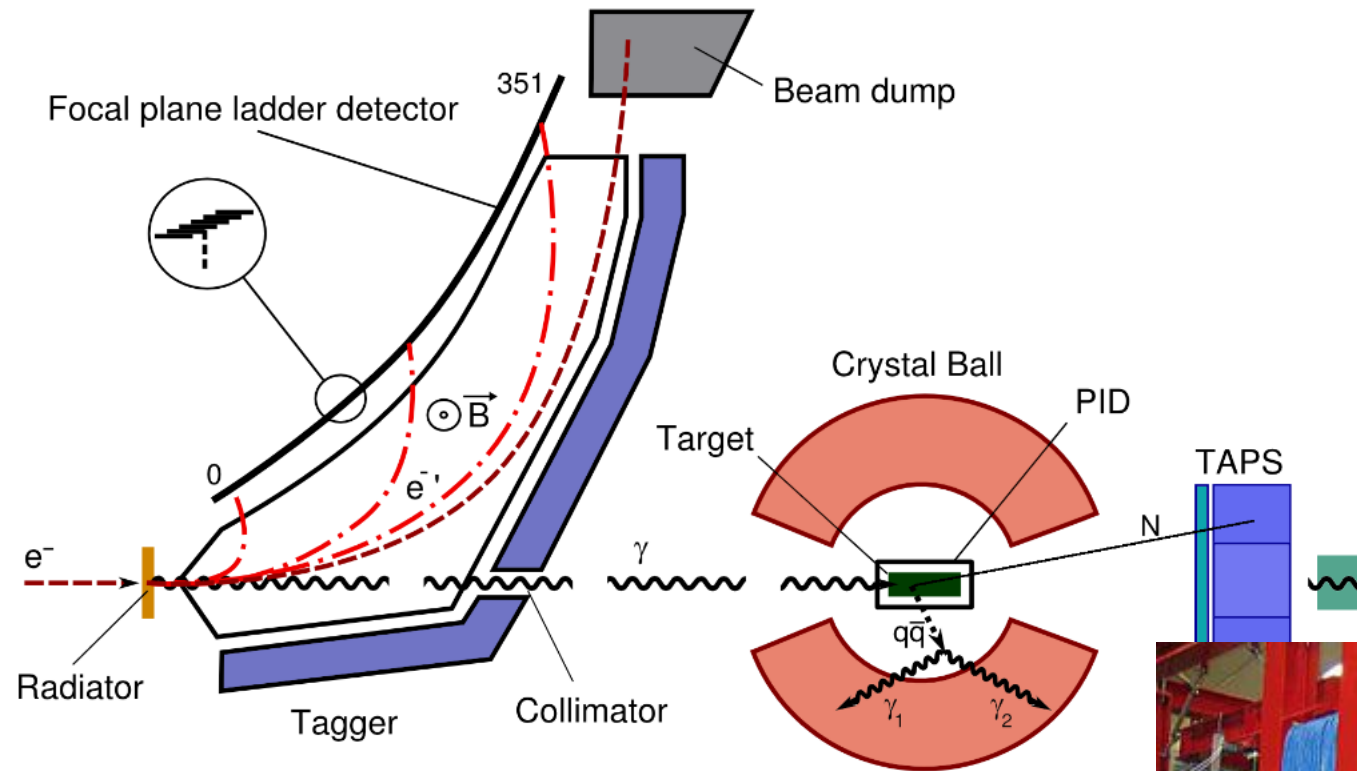
Scintillation efficiency = 1 photon/100 eV

Collection efficiency = 0.10% (large counter)

Quantum efficiency = 0.25%

$$Y = Y_{lab} \cdot (1 - a) \cdot \epsilon_{cov} \cdot \epsilon_{PMT} = 10^4 / \text{MeV} \cdot 0.42 \cdot 0.75 \cdot 0.25 = 1100 / \text{MeV}$$

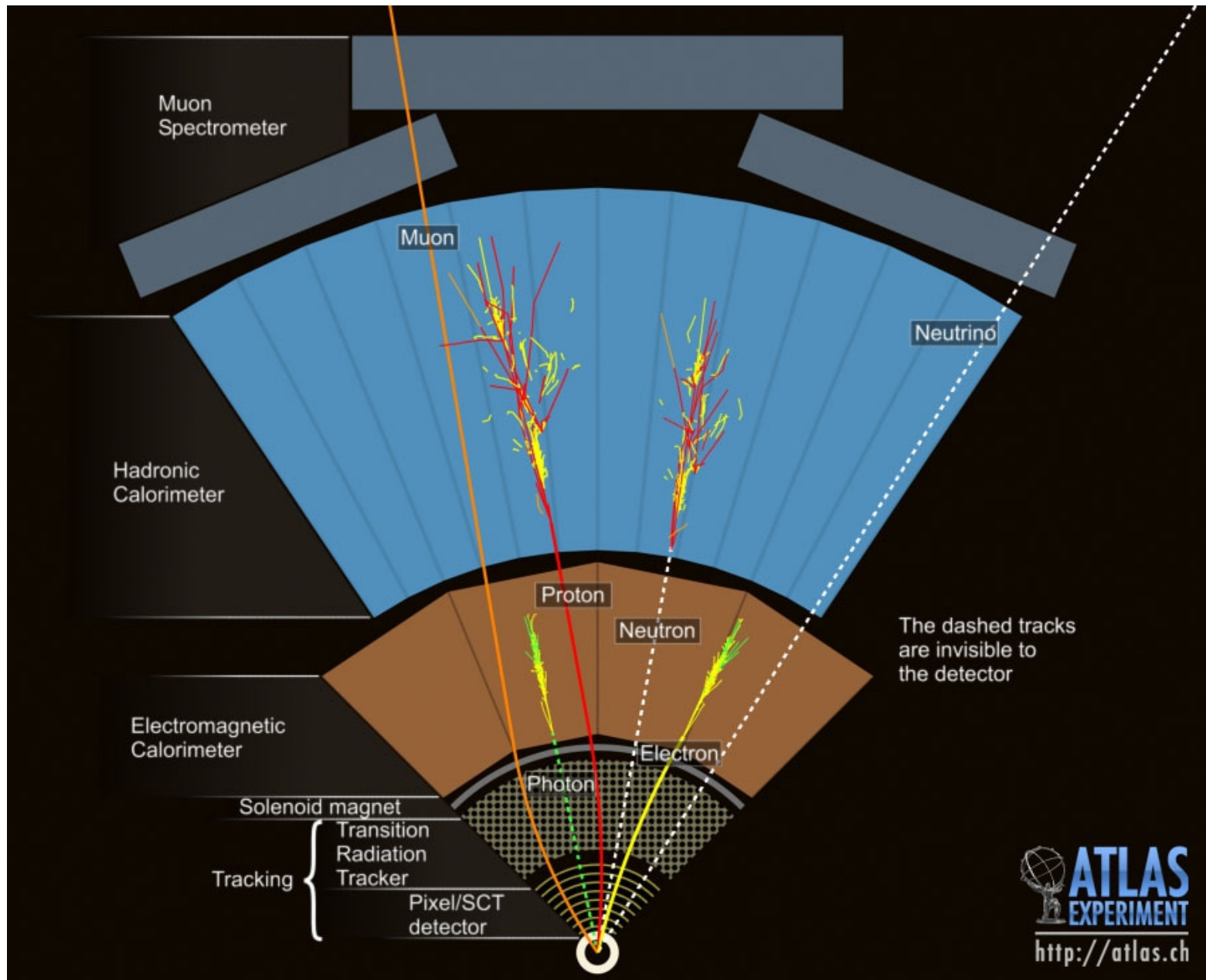
$$\Delta E / E = \Delta Y / Y \approx \sqrt{Y} / Y \approx 0.03$$

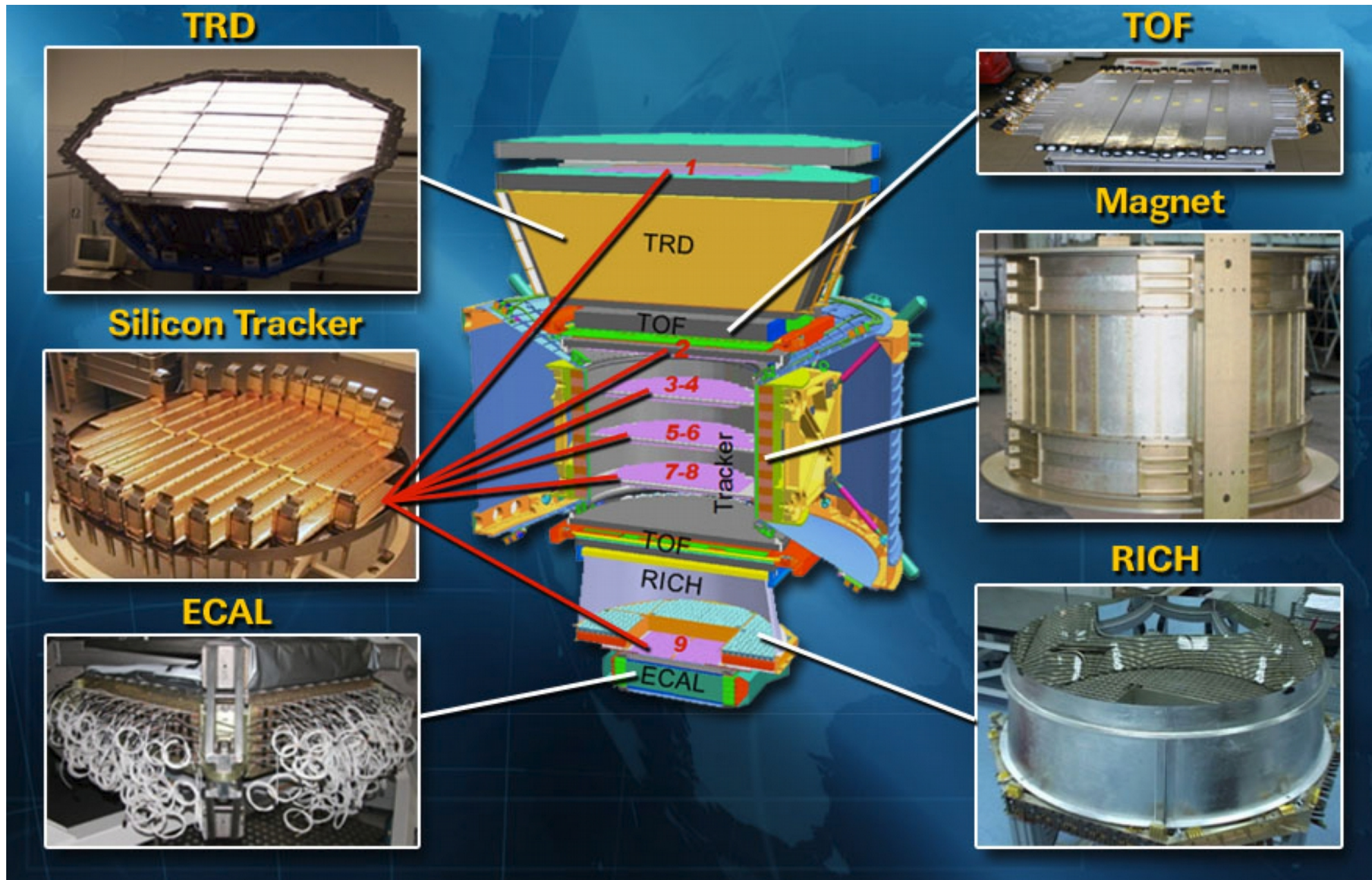


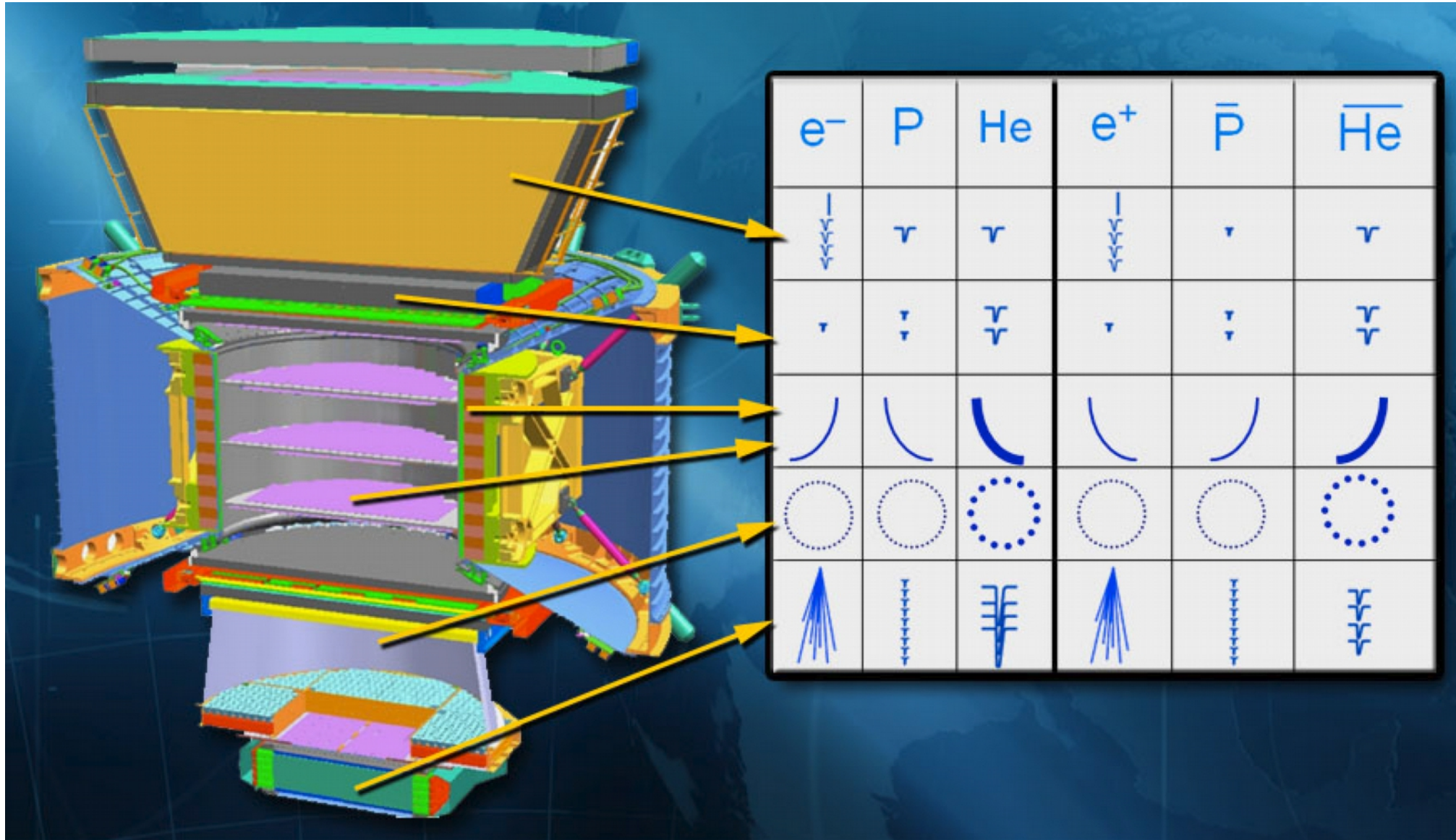
Commonly Used Inorganic Scintillators

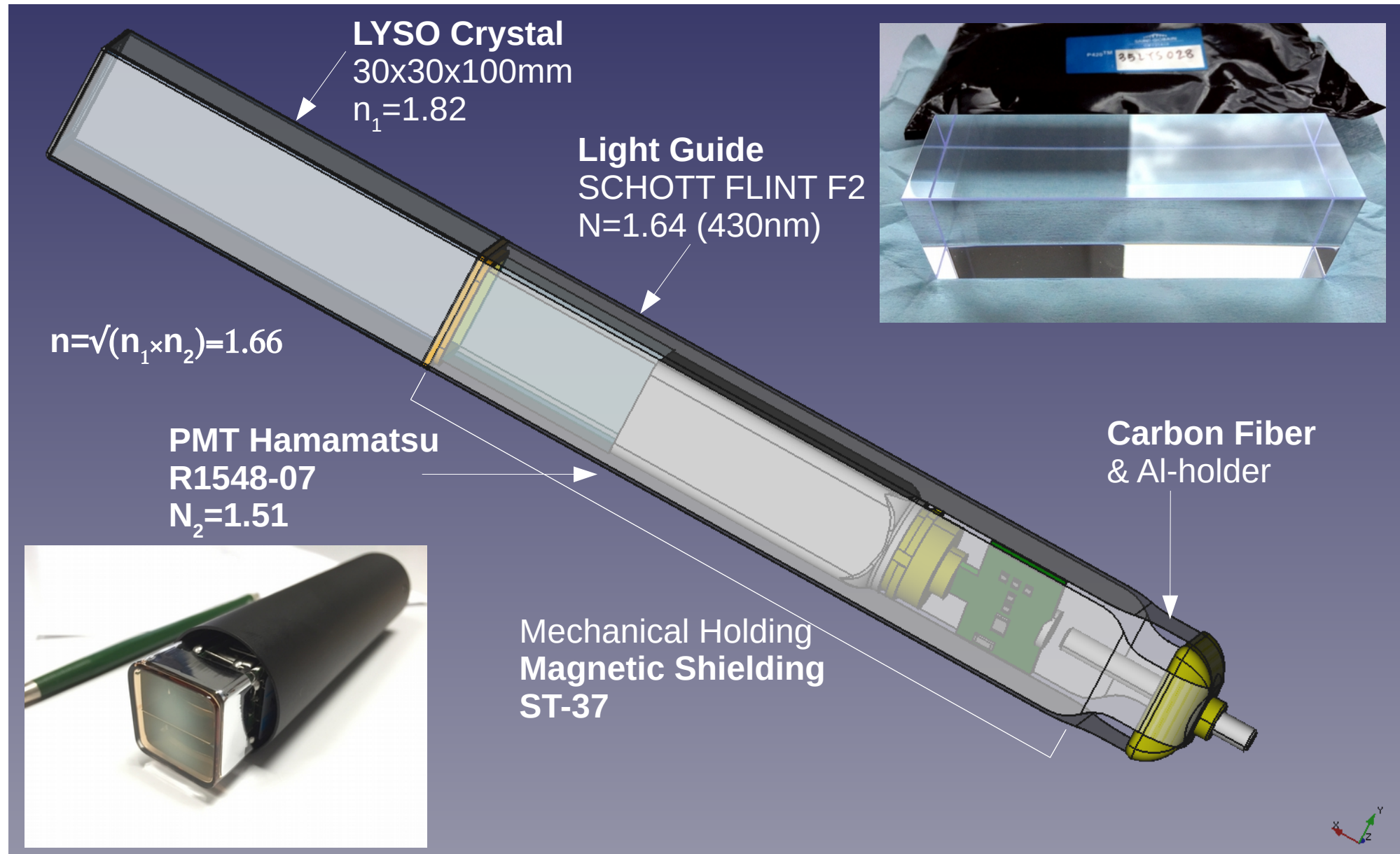
Scintillator	Light yield (photons/keV)	Light output (%) of NaI(Tl) bialkali pmt	Temperature coefficient of light output (%/°C) 25°C to 50°C	1/e Decay time (ns) (10-3 μs)	Wavelength of maximum emission γm (nm)	Refractive index at γm	Thickness to stop 50% of 662 keV photons (cm)	Thermal expansion (°/C) x 10 ⁶	Hardness (Mho)	Density g/cm ³	Hygroscopic	Comments
BrilLanCe™380 LaBr ₃ (Ce)	63	165	0	16	380	~1.9	1.8	8		5.08	yes	General purpose, best energy resolution, rate of change of light output w/ temperature is small
NaI(Tl)	38	100	-0.3	250	415	1.85	2.5	47.4	2	3.67	yes	General purpose, good energy resolution
Polyscin®NaI(Tl)	38	100	-0.3	250	415	1.85	2.5	47.4	2	3.67	yes	Polycrystalline NaI(Tl), for extra strength
BrilLanCe™350 LaCl ₃ (Ce)	49	70 - 90	0.7	28	350	~1.9	2.3	11		3.85	yes	General purpose, excellent energy resolution
CsI(Na)	41	85	-0.05	630	420	1.84	2	54	2	4.51	yes	High Z, rugged
PreLude™420 Lu _{1.82} Y _{0.18} SiO ₅ (Ce)	32	75	-0.28	41	420	1.81	1.1	--		7.1	no	Bright, high Z, fast, dense, background from ¹⁷⁶ Lu activity
CdWO ₄	12 - 15	30 - 50	-0.1	14000	475	~2.3	1	10.2	4 - 4.5	7.9	no	High Z, low afterglow, for use with photodiodes
CaF ₂ (Eu)	19	50	-0.33	940	435	1.47	2.9	19.5	4	3.18	no	Low Z, α & β detection
CsI(Tl)	54	45	0.01	1000	550	1.79	2	54	2	4.51	slightly	High Z, rugged, good match to photodiodes
BGO	8 - 10	20	-1.2	300	480	2.15	1	7	5	7.13	no	High Z, compact detector, low afterglow
YAG(Ce) Y ₃ Al ₅ O ₁₂ (Ce)	8	15	--	70	550	1.82	2	~80	8.5	4.55	no	β-ray, X-ray counting, electron microscopy
CsI(pure)	2	4 - 6	-0.3	16	315	1.95	2	54	2	4.51	slightly	High Z, fast emission
BaF ₂	1.8	3	0	0.6 - 0.8	220 (195)	1.54	1.9	18.4	3	4.88	slightly	Fast component (subnanosecond)
	10	16	-1.1	630	310	1.50	1.9	18.4	3	4.88	slightly	Slow component
ZnS(Ag)	~50	130	-0.6	110	450	2.36	--	--	--	4.09	no	Multicrystal, 15μ stops 5.5 MeV α (n detection with ⁶ Li)

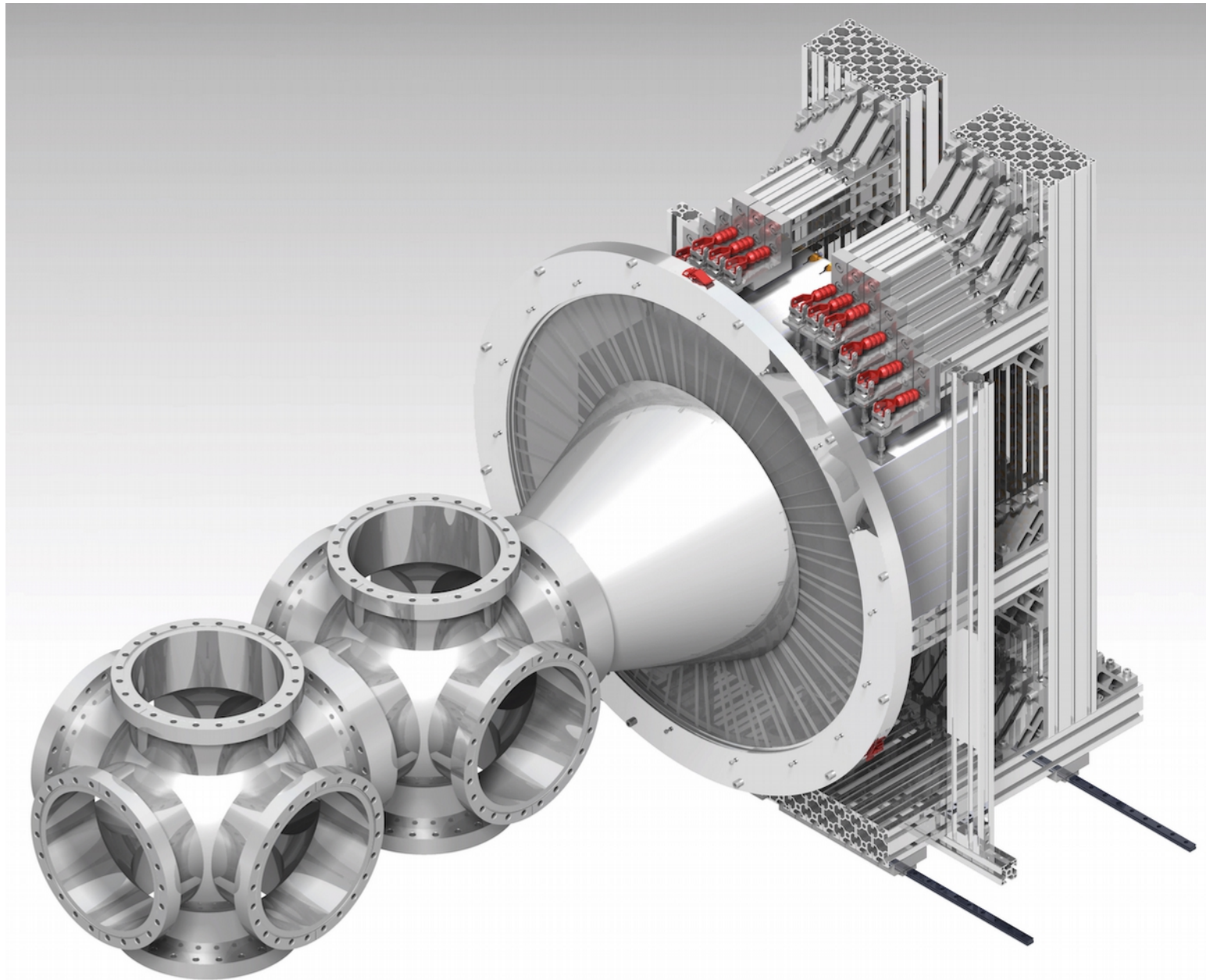
The data presented are believed to be correct but are not guaranteed to be so.

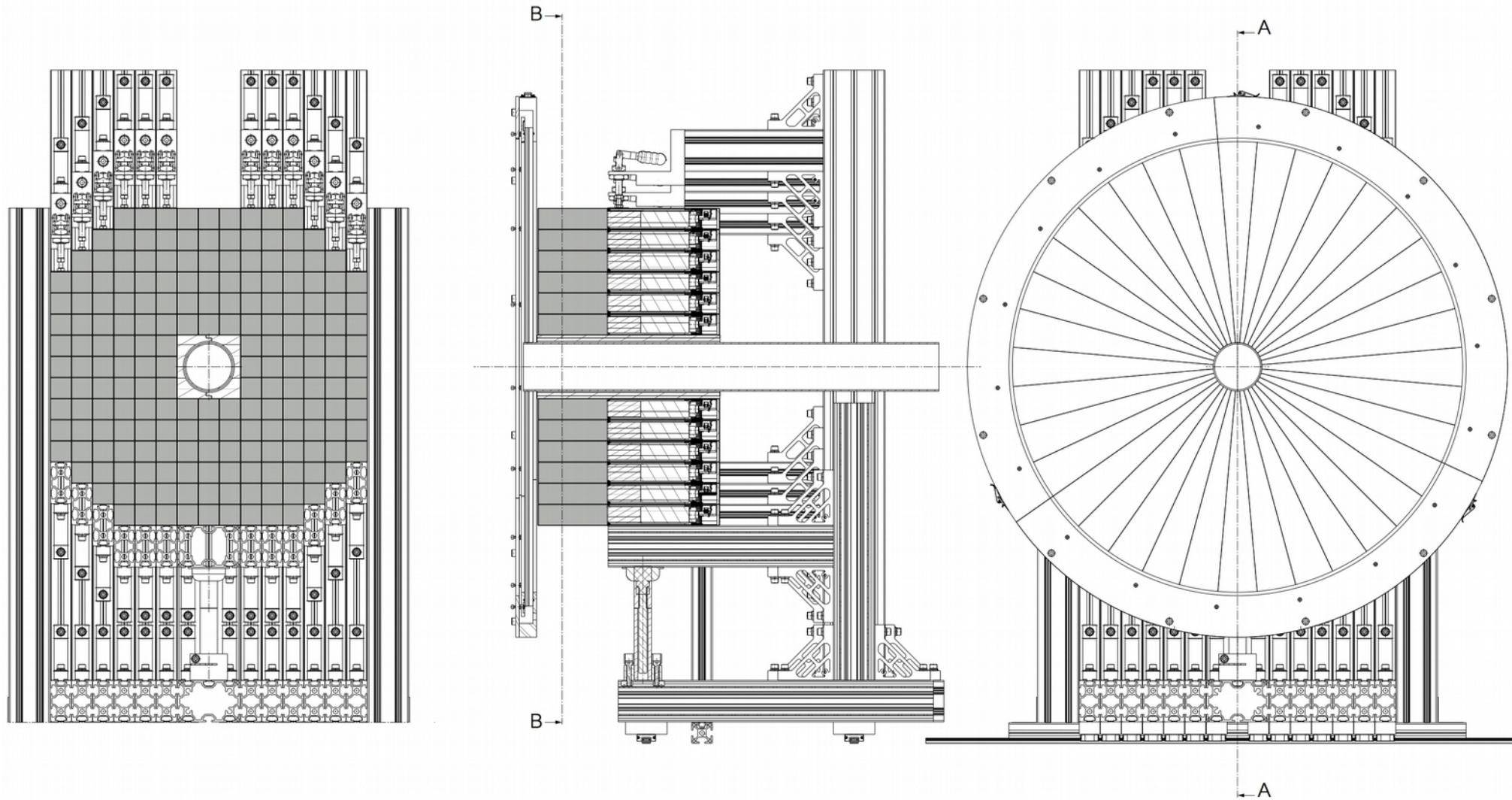












- Purpose of the Detectors
- Working Principle of the Detectors
- How to choose the Detector Types/Materials
- Some Examples